Cold atoms

Lecture 5. 15th November, 2006

Preliminary plan/reality in the fall term

Lecture 1	Something about everything (see next slide) The textbook version of BEC in extended systems	Sep 22
Lecture 2	thermodynamics, grand canonical ensemble, extended gas: ODLRO, nature of the BE phase transition	Oct 4
Lecture 3	atomic clouds in the traps – independent bosons, what is BEC?, "thermodynamic limit", properties of OPDM	Oct 18
Lecture 4	atomic clouds in the traps – interactions, GP equation at zero temperature, variational prop., chem. potential	Nov 1
Lecture 5	Infinite systems: Bogolyubov theory, BEC and symmetry breaking, coherent states	Nov 15
Lecture 6	Time dependent GP theory. Finite systems: BEC theory preserving the particle number	

Previous class: Interacting atoms

L4: Scattering length, pseudopotential

Beyond the potential radius, say 3σ , the scattered wave propagates in free space

For small energies, the scattering is purely isotropic, the *s*-wave scattering. The outside wave is

$$\psi \propto \frac{\sin(kr + \delta_0)}{r}$$

For very small energies the radial part becomes just

$$r - a_s$$
, a_s ... the scattering length

This may be extrapolated also into the interaction sphere (we are not interested in the short range details)

Equivalent potential ("pseudopotential")

$$U(r) = g \cdot \delta(r)$$
$$g = \frac{4\pi a_s \hbar^2}{m}$$

Previous class: Mean-field treatment of interacting atoms L4: Many-body Hamiltonian and the Hartree approximation

$$\hat{H} = \sum_{a} \frac{1}{2m} p_{a}^{2} + V(\mathbf{r}_{a}) + \frac{1}{2} \sum_{a \neq b} \sum_{b} U(\mathbf{r}_{a} - \mathbf{r}_{b})$$

We start from the mean field approximation.

This is an educated way, similar to (almost identical with) the HARTREE APPROXIMATION we know for many electron systems.

Most of the interactions is indeed absorbed into the mean field and what remains are explicit quantum correlation corrections

$$\hat{H}_{\rm GP} = \sum_{a} \frac{1}{2m} p_a^2 + V(\mathbf{r}_a) + V_H(\mathbf{r}_a)$$

$$V_H(\mathbf{r}_a) = \int d\mathbf{r}_b U(\mathbf{r}_a - \mathbf{r}_b) n(\mathbf{r}_b) = g \cdot n(\mathbf{r}_a)$$

$$n(\mathbf{r}) = \sum_{\alpha} n_{\alpha} \left| \varphi_{\alpha} \left(\mathbf{r} \right) \right|^2$$

$$\left(\frac{1}{2m} p^2 + V(\mathbf{r}) + V_H(\mathbf{r}) \right) \varphi_{\alpha} \left(\mathbf{r} \right) = E_{\alpha} \varphi_{\alpha} \left(\mathbf{r} \right)$$

L4: Gross-Pitaevskii equation at zero temperature

Consider a condensate. Then all occupied orbitals are the same and we have a single self-consistent equation for a single orbital

$$\left(\frac{1}{2m}p^{2}+V(\boldsymbol{r})+gN\left|\boldsymbol{\varphi}_{0}\left(\boldsymbol{r}\right)\right|^{2}\right)\boldsymbol{\varphi}_{0}\left(\boldsymbol{r}\right)=E_{0}\boldsymbol{\varphi}_{0}\left(\boldsymbol{r}\right)$$

Putting

$$\Psi(\boldsymbol{r}) = \sqrt{N} \cdot \boldsymbol{\varphi}_0(\boldsymbol{r})$$

we obtain a closed equation for the order parameter:

The lowest level coincides with the chemical potential

$$\left(\frac{1}{2m}p^2 + V(\mathbf{r}) + g\left|\boldsymbol{\Psi}(\mathbf{r})\right|^2\right)\boldsymbol{\Psi}(\mathbf{r}) = \boldsymbol{\mu}\boldsymbol{\Psi}(\mathbf{r})$$

This is the celebrated Gross-Pitaevskii equation.

For a static condensate, the order parameter has ZERO PHASE. Then $\Psi(\mathbf{r}) = \sqrt{N} \cdot \varphi_0(\mathbf{r}) = \sqrt{n(\mathbf{r})}$

$$N[n] = N = \int d^3 \mathbf{r} |\Psi(\mathbf{r})|^2 = \int d^3 \mathbf{r} \cdot n(\mathbf{r}) = N$$

Gross-Pitaevskii equation – homogeneous gas

The GP equation simplifies

$$\left(-\frac{\hbar^2}{2m}\Delta + g\left|\Psi(\mathbf{r})\right|^2\right)\Psi(\mathbf{r}) = \mu\Psi(\mathbf{r})$$

For periodic boundary conditions in a box with $V = L_x \cdot L_y \cdot L_z$

$$\varphi_{0}(\mathbf{r}) = \frac{1}{\sqrt{V}}$$

$$\Psi(\mathbf{r}) = \sqrt{N} \cdot \varphi_{0}(\mathbf{r}) = \sqrt{\frac{N}{V}} = \sqrt{n}$$

$$g |\Psi(\mathbf{r})|^{2} \Psi(\mathbf{r}) = \mu \Psi(\mathbf{r}) \quad \dots \text{ GP equation}$$

$$|\underline{\mu} = g |\Psi(\mathbf{r})|^{2} = \underline{gn}$$

$$|\underline{E}_{\underline{N}} = \frac{1}{N} \int d^{3}\mathbf{r} \left\{ \frac{\hbar^{2}}{2m} (\nabla \sqrt{n})^{2} + \overline{V}(\mathbf{r})n + \frac{1}{2}gn^{2} \right\} = \frac{1}{2}gn$$

Field theoretic reformulation (second quantization)

Field operator for spin-less bosons

Definition by commutation relations

$$\left[\psi(\mathbf{r}),\psi^{\dagger}(\mathbf{r'})\right] = \delta(\mathbf{r}-\mathbf{r'}), \quad \left[\psi(\mathbf{r}),\psi(\mathbf{r'})\right] = 0, \quad \left[\psi^{\dagger}(\mathbf{r}),\psi^{\dagger}(\mathbf{r'})\right] = 0$$

basis of single-particle states (κ complete set of quantum numbers)

$$\{ |\kappa\rangle \} \quad \langle \kappa |\beta\rangle = \delta_{\kappa\beta} \quad |\psi\rangle = \sum |\kappa\rangle \langle \kappa |\psi\rangle, \quad \psi \quad ... \text{ single particle state} \langle r |\kappa\rangle = \varphi_{\kappa}(r) \quad \langle r |\psi\rangle = \sum \langle r |\kappa\rangle \langle \kappa |\psi\rangle$$

decomposition of the field operator

$$\psi(\mathbf{r}) = \sum \varphi_{\kappa}(\mathbf{r}) a_{\kappa}, \quad a_{\kappa} = "\langle \kappa | \psi \rangle" = \int d^{3} \varphi_{\kappa}^{*}(\mathbf{r}) \psi(\mathbf{r})$$

$$\psi^{\dagger}(\mathbf{r}) = \sum \varphi_{\kappa}^{*}(\mathbf{r}) a_{\kappa}^{\dagger}$$

commutation relations

$$\begin{bmatrix} a_{\kappa}, a_{\lambda}^{\dagger} \end{bmatrix} = \delta_{\kappa\lambda}, \quad \begin{bmatrix} a_{\kappa}, a_{\lambda} \end{bmatrix} = 0, \quad \begin{bmatrix} a_{\kappa}^{\dagger}, a_{\lambda}^{\dagger} \end{bmatrix} = 0$$

Field operator for spin-less bosons – cont'd

Plane wave representation (BK normalization)

$$\psi(\mathbf{r}) = V^{-1/2} \sum e^{i\mathbf{k}\mathbf{r}} a_k, \quad a_k = V^{-1/2} \int d^3 \mathbf{r} e^{-i\mathbf{k}\mathbf{r}} \psi(\mathbf{r})$$

$$\psi^{\dagger}(\mathbf{r}) = V^{-1/2} \sum e^{-i\mathbf{k}\mathbf{r}} a_k^{\dagger} = V^{-1/2} \sum e^{i\mathbf{k}\mathbf{r}} a_{-\mathbf{k}}^{\dagger}$$

$$\begin{bmatrix} a_k, a_{k'}^{\dagger} \end{bmatrix} = \delta_{kk'}, \quad \begin{bmatrix} a_k, a_{k'} \end{bmatrix} = 0, \quad \begin{bmatrix} a_k^{\dagger}, a_{k'}^{\dagger} \end{bmatrix} = 0$$

Operators

Additive observable

$$\boldsymbol{X} = \sum X_{j} \quad \rightarrow \quad \boldsymbol{X} = \iint d^{3}\boldsymbol{r} \, d^{3}\boldsymbol{r'} \, \boldsymbol{\psi}^{\dagger}(\boldsymbol{r}) \left\langle \boldsymbol{r} \right| \boldsymbol{X} \left| \boldsymbol{r'} \right\rangle \boldsymbol{\psi}(\boldsymbol{r'})$$

General definition of the OPDM

Particle number

$$N = \sum 1_{\text{OP},j} \quad \rightarrow \quad N = \int d^3 r \, \psi^{\dagger}(r) \psi(r)$$
$$N = \sum a_{\kappa}^{\dagger} a_{\kappa}$$

Hamiltonian

$$H = \sum_{a} \frac{1}{2m} p_{a}^{2} + V(\mathbf{r}_{a}) + \frac{1}{2} \sum_{a \neq b} U(\mathbf{r}_{a} - \mathbf{r}_{b})$$

= $\int d^{3}\mathbf{r} \psi^{\dagger}(\mathbf{r}) \Big(-\frac{\hbar^{2}}{2m} \Delta + V(\mathbf{r}) \Big) \psi(\mathbf{r}) + \frac{1}{2} \iint d^{3}\mathbf{r} d^{3}\mathbf{r'} \psi^{\dagger}(\mathbf{r}) \psi^{\dagger}(\mathbf{r'}) U(\mathbf{r} - \mathbf{r'}) \psi(\mathbf{r'}) \psi(\mathbf{r'})$

Particle number conservation

$$[H, N] = 0$$

Equilibrium density operators and the ground state

$$P = P(H), [N, p] = 0$$

Typical selection rule

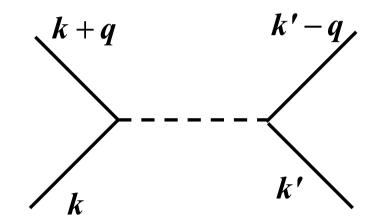
$$\langle \psi(\mathbf{r}) \rangle = \operatorname{Tr} \psi(\mathbf{r}) \mathbf{p} = 0$$

is a consequence of the gauge invariance of the 1st kind:

$$\operatorname{Tr} \psi P = \operatorname{Tr} \psi e^{i\varphi N} P e^{-i\varphi N} = \operatorname{Tr} e^{-i\varphi N} \psi e^{i\varphi N} P = e^{-i\varphi} \operatorname{Tr} \psi P$$

Hamiltonian of the homogeneous gas

$$H = \sum \frac{\hbar^2}{2m} k^2 a_k^{\dagger} a_k + \frac{1}{2} V^{-1} \sum_{kk'q} U_q a_{k+q}^{\dagger}, a_{k'-q}^{\dagger} a_{k'} a_k,$$
$$U_k = \int d^3 \mathbf{r} e^{-i\mathbf{k}\mathbf{r}} U(\mathbf{r})$$



Action of the field operators in the Fock space

basis of single-particle states $\{|\kappa\rangle\} \quad \langle\kappa|\beta\rangle = \delta_{\kappa\beta} \quad |\psi\rangle = \sum |\kappa\rangle\langle\kappa|\psi\rangle, \quad \psi \quad ... \text{ single particle state}$ $\langle r|\kappa\rangle = \varphi_{\kappa}(r) \quad \langle r|\psi\rangle = \sum \langle r|\kappa\rangle\langle\kappa|\psi\rangle$

FOCK SPACE space of many particle states

basis states ... symmetrized products of single-particle states for bosons specified by the set of **occupation numbers** 0, 1, 2, 3, ... $\{\kappa_1, \kappa_2, \kappa_3, ..., \kappa_p, ...\}$ $\Psi_{\{n_\kappa\}} = | n_1, n_2, n_3, ..., n_p, ... \rangle$ *n*-particle state $n = \Sigma n_p$ $a_p^{\dagger} | n_1, n_2, n_3, ..., n_p, ... \rangle = \sqrt{n_p + 1} | n_1, n_2, n_3, ..., n_p + 1, ... \rangle$ $a_p | n_1, n_2, n_3, ..., n_p, ... \rangle = \sqrt{n_p} | n_1, n_2, n_3, ..., n_p - 1, ... \rangle$

Bogolyubov method

Basic idea

Bogolyubov method

is devised for boson quantum fluids with weak interactions – at T=0 now

no interaction		weak interaction	
g = 0		$g \neq 0$	
$N = N_{\rm BE} = \left\langle a_0^{\dagger} a_0 \right\rangle$	1	$N = N_{\rm BE} + \sum_{k \neq 0} \left\langle a_k^{\dagger} a_k \right\rangle \approx N_{\rm BE}$	1

The condensate dominates.

Strange idea

$$N_{0} = \left\langle a_{0}^{\dagger} a_{0} \right\rangle \quad 1 \Longrightarrow \left\langle a_{0}^{\dagger} a_{0} \right\rangle \quad a_{0}^{\dagger} a_{0} - a_{0} a_{0}^{\dagger} \Longrightarrow \text{like } c\text{-numbers}$$
$$\boxed{a_{0} \approx \sqrt{N_{0}}, \quad a_{0}^{\dagger} \approx \sqrt{N_{0}}}$$
$$N = N_{0} + \sum_{k \neq 0} a_{k}^{\dagger} a_{k} \quad \dots \text{ mixture of } c\text{-numbers and } q\text{-numbers}$$

Approximate Hamiltonian

Keep at most two particles out of the condensate

$$H = \sum \frac{\hbar^2}{2m} k^2 a_k^{\dagger} a_k + \frac{1}{2} V^{-1} \sum_{kk'q} U_q a_{k+q}^{\dagger}, a_{k'-q}^{\dagger} a_{k'} a_k$$

$$= \sum \frac{\hbar^2}{2m} k^2 a_k^{\dagger} a_k + \frac{UN_0}{2V} \sum_k \left\{ a_k^{\dagger} a_{-k}^{\dagger} + 4a_k^{\dagger} a_k + a_k a_{-k} \right\} + \frac{UN_0^2}{2V}$$

$$= \sum \frac{\hbar^2}{2m} k^2 a_k^{\dagger} a_k + \frac{UN}{2V} \sum_k \left\{ a_k^{\dagger} a_{-k}^{\dagger} + 2a_k^{\dagger} a_k + a_k a_{-k} \right\} + \frac{UN^2}{2V}$$

Bogolyubov transformation

Last rearrangement

$$H = \frac{1}{2} \sum_{k} \underbrace{\left(\frac{\hbar^2}{2m} k^2 + gn\right)}_{\text{mean field}} \left\{ a_k^{\dagger} a_k + a_{-k}^{\dagger} a_{-k} \right\} + \frac{gn}{2} \sum_{k} \underbrace{\left\{ a_k^{\dagger} a_{-k}^{\dagger} + a_k a_{-k} \right\}}_{\text{anomalous}} + \frac{UN^2}{2V}$$

Conservation properties: momentum ... YES, particle number ... NO
<u>NEW FIELD OPERATORS</u> notice momentum conservation!!

requirements

- New operators should satisfy the boson commutation rules $\begin{bmatrix} b_k, b_{k'}^{\dagger} \end{bmatrix} = \delta_{kk'}, \quad \begin{bmatrix} b_k, b_{k'} \end{bmatrix} = 0, \quad \begin{bmatrix} b_k^{\dagger}, b_{k'}^{\dagger} \end{bmatrix} = 0$ iff $u_k^2 - v_k^2 = 1$
- When introduced into the Hamiltonian, the anomalous terms have to vanish

Bogolyubov transformation – result

Without quoting the transformation matrix

H =
$$\frac{1}{2}\sum_{\substack{\varepsilon(k) \{b_k^{\dagger}b_k + b_{-k}^{\dagger}b_{-k}\} \\ \text{independent quasiparticles}}} + \frac{UN^2}{2V} + \text{higher order constant}}$$

 $\varepsilon(k) = \sqrt{\left(\frac{\hbar^2}{2m}k^2 + gn\right)^2 - (gn)^2} = \sqrt{\frac{\hbar^2}{2m}k^2}\sqrt{\frac{\hbar^2}{2m}k^2 + 2gn}$
high energy region
quasi-particles are
nearly just particles

$$\varepsilon(k) \qquad \qquad \frac{\hbar^2}{2m}k^2 \qquad \qquad \text{merge} \\ \omega(k) = c \cdot k \\ k \rightarrow \qquad \qquad c = \sqrt{\frac{gn}{m}} \\ \end{array}$$

 \uparrow

sound region

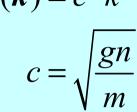
quasi-particles are collective excitations

More about the sound part of the dispersion law

Entirely dependent on the interactions, both the magnitude of the velocity and the linear frequency range determined by g $\omega(k) = c \cdot k$

Can be shown to really be a sound:

$$c = "\sqrt{\frac{\kappa}{\rho}}" = \sqrt{\frac{V\partial_{VV}E}{m \cdot n}}, \qquad E = \frac{UN^2}{2V} + \cdots$$



Even a weakly interacting gas exhibits superfluidity; the ideal gas does not.

- The phonons are actually Goldstone modes corresponding to a broken symmetry
- The dispersion law has no roton region, contrary to the reality
- The dispersion law bends upwards ⇒ quasi-particles are unstable, can decay

Particles and quasi-particles

At zero temperature, there are no quasi-particles, just the condensate.

Things are different with the true particles. Not <u>all</u> particles are in the condensate, but they are not thermally agitated in an incoherent way, they are a part of the fully coherent ground state

$$\left\langle a_{k}^{\dagger}a_{k}^{\dagger}\right\rangle = \left\langle \left(-v_{k}^{\dagger}b_{k}^{\dagger}+u_{k}^{\dagger}b_{-k}^{\dagger}\right)\left(u_{k}^{\dagger}b_{-k}^{\dagger}-v_{k}^{\dagger}b_{k}^{\dagger}\right)\right\rangle = v_{k}^{2} \neq 0$$

The total amount of the particles outside of the condensate is

$$\frac{N - N_0}{N} \approx \frac{8}{3\sqrt{\pi}} \underbrace{a_s^{3/2} n^{1/2}}_{\sqrt{a_s^3 n}} \qquad \text{the gas parameter} \\ \text{is} \\ \text{the expansion variable} \end{cases}$$

Coherent ground state

Reformulation of the Bogolyubov requirements

Looks like he wanted

$$a_0 |\Psi\rangle = \Psi |\Psi\rangle, \quad \Psi = \sqrt{N}, \text{ so that}$$

 $\langle a_0 \rangle = \Psi$

This is in contradiction with the rule derived above, $\langle a_0 \rangle = 0$

The above equation is known and defines the ground state to be a coherent state with the parameter Ψ

For a coherent state, there is no problem with the particle number conservation. It has a rather uncertain particle number, but a well defined phase:

$$|\Psi\rangle = |\Psi\rangle = e^{-|\Psi|^{2}/2} \cdot e^{\Psi a_{0}^{\dagger}} |\operatorname{vac}\rangle$$
$$\langle\Psi|a_{0}|\Psi\rangle = \Psi$$
$$\langle\Psi|a_{0}^{\dagger}a_{0}|\Psi\rangle = |\Psi|^{2}$$
$$\langle\Psi|a_{0}^{\dagger}a_{0}a_{0}^{\dagger}a_{0}|\Psi\rangle = |\Psi|^{4} + |\Psi|^{2}$$
$$\Delta N_{0} = |\Psi|$$

The end

On the way to the mean-field Hamiltonian

ADDITIONAL NOTES On the way to the mean-field Hamiltonian

① First, the following <u>exact</u> transformations are performed

$$\hat{H} = \sum_{a} \frac{1}{2m} p_{a}^{2} + \sum_{a} V(r_{a}) + \frac{1}{2} \sum_{a \neq b} U(r_{a} - r_{b})$$

$$\hat{V} = \sum_{a} V(r_{a}) = \int d^{3}r V(r) \sum_{a} \delta(r - r_{a}) \equiv \int d^{3}r V(r) \cdot \hat{n}(r)$$

$$particle density operator$$

$$\hat{U} = \frac{1}{2} \sum_{a \neq b} U(r_{a} - r_{b}) = \frac{1}{2} \int d^{3}r d^{3}r' U(r - r') \sum_{a \neq b} \delta(r - r_{a}) \delta(r' - r_{b})$$

$$= \frac{1}{2} \int d^{3}r d^{3}r' U(r - r') \sum_{a} \delta(r - r_{a}) \left\{ \sum_{b} \delta(r' - r_{b}) - \frac{\nabla(r' - r')}{\delta(r' - r')} \right\}$$

$$\hat{H} = \hat{W} + \int d^{3}r V(r) \cdot \hat{n}(r) + \frac{1}{2} \int d^{3}r d^{3}r' U(r - r') \hat{n}(r) \left\{ \hat{n}(r') - \delta(r - r') \right\}$$

$$27$$

On the way to the mean-field Hamiltonian

② Second, a specific many-body state is chosen, which defines the mean field:

$$\Psi \to n(\mathbf{r}) = \left\langle \hat{n}(\mathbf{r}) \right\rangle \equiv \left\langle \Psi \right| \hat{n}(\mathbf{r}) \left| \Psi \right\rangle$$

Then, the operator of the (quantum) density fluctuation is defined:

$$\hat{n}(\mathbf{r}) = n(\mathbf{r}) + \Delta \hat{n}(\mathbf{r})$$

$$\hat{n}(\boldsymbol{r})\hat{n}(\boldsymbol{r'}) = \hat{n}(\boldsymbol{r})n(\boldsymbol{r'}) + n(\boldsymbol{r})\hat{n}(\boldsymbol{r'}) + \Delta\hat{n}(\boldsymbol{r})\Delta\hat{n}(\boldsymbol{r'}) - n(\boldsymbol{r})n(\boldsymbol{r'})$$

The Hamiltonian, still exactly, becomes

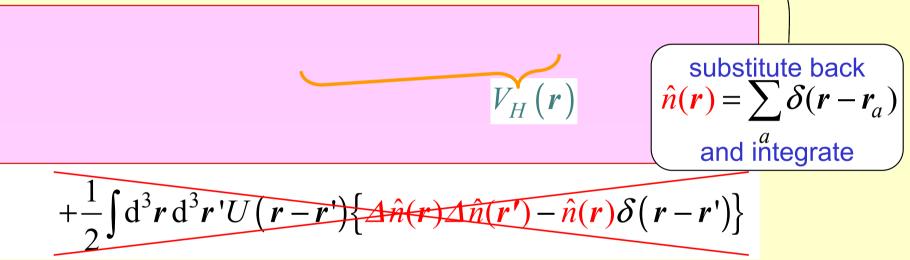
$$\hat{H} = \hat{W} + \int d^3r \left\{ V(r) + \int d^3r' U(r-r')n(r') \right\} \cdot \hat{n}(r)$$

$$-\frac{1}{2} \int d^3r d^3r' U(r-r')n(r)n(r')$$

$$+\frac{1}{2} \int d^3r d^3r' U(r-r') \left\{ \Delta \hat{n}(r) \Delta \hat{n}(r') - \hat{n}(r) \delta(r-r') \right\}$$

ADDITIONAL NOTES On the way to the mean-field Hamiltonian

③ In the last step, the third line containing exchange, correlation and the self-interaction correction is neglected. The mean-field Hamiltonian of the main lecture results:



REMARKS

• Second line ... an additive constant compensation for doublecounting of the Hartree interaction energy

• In the original (variational) Hartree approximation, the self-interaction is not left out, leading to non-orthogonal Hartree orbitals

BACK

Variational approach to the condensate ground state

Variational estimate of the condensate properties

VARIATIONAL PRINCIPLE OF QUANTUM MECHANICS
 The ground state and energy are uniquely defined by

$$E = \left\langle \Psi \middle| \hat{H} \middle| \Psi \right\rangle \leq \left\langle \Psi' \middle| \hat{H} \middle| \Psi' \right\rangle \quad \text{for all } \left| \Psi' \right\rangle \in \mathcal{H}_N^S, \left\langle \Psi' \middle| \Psi' \right\rangle = 1$$

In words, $|\Psi'\rangle$ is a normalized symmetrical wave function of *N* particles. The minimum condition in the variational form is

 $\delta \langle \Psi | \hat{H} | \Psi \rangle = 0$ equivalent with the SR $\hat{H} | \Psi \rangle = E | \Psi \rangle$

ARTREE VARIATIONAL ANSATZ FOR THE CONDENSATE WAVE F.
 For our many-particle Hamiltonian,

$$\hat{H} = \sum_{a} \frac{1}{2m} p_{a}^{2} + V(\mathbf{r}_{a}) + \frac{1}{2} \sum_{a \neq b} \sum_{b} U(\mathbf{r}_{a} - \mathbf{r}_{b}), \quad U(\mathbf{r}) = g \cdot \delta(\mathbf{r})$$

the true ground state is approximated by the condensate for non-interacting particles (Hartree Ansatz, here identical with the symmetrized Hartree-Fock)

$$\Psi(\mathbf{r}_1,\mathbf{r}_2,\ldots,\mathbf{r}_p,\ldots,\mathbf{r}_N) = \varphi_0(\mathbf{r}_1)\varphi_0(\mathbf{r}_2)\cdots\varphi_0(\mathbf{r}_p)\cdots\varphi_0(\mathbf{r}_N)$$

Variational estimate of the condensate properties

Here, φ_0 is a normalized real spinless orbital. It is a functional variable to be found from the variational condition

$$\delta \mathcal{E}[\varphi_0] = \delta \langle \Psi[\varphi_0] | \hat{H} | \Psi[\varphi_0] \rangle = 0 \quad \text{with} \quad \langle \Psi[\varphi_0] | \Psi[\varphi_0] \rangle = 1 \Leftrightarrow \langle \varphi_0 | \varphi_0 \rangle = 1$$

Explicit calculation yields

$$\mathcal{E}[\varphi_0] = \frac{\hbar^2}{2m} N \int \mathrm{d}^3 \mathbf{r} \left(\nabla \varphi_0(\mathbf{r}) \right)^2 + N \int \mathrm{d}^3 \mathbf{r} V(\mathbf{r}) \left(\varphi_0(\mathbf{r}) \right)^2 + \frac{1}{2} N \left(N - 1 \right) g \int \mathrm{d}^3 \mathbf{r} \left(\varphi_0(\mathbf{r}) \right)^4$$

Variation of energy with the use of a Lagrange multiplier:

$$\delta \Big\{ N^{-1} \mathcal{E}[\varphi_0] - \mu \big\langle \varphi_0 \big| \varphi_0 \big\rangle \Big\} \qquad \qquad \varphi_0 = \varphi_0(\mathbf{r}), \ \delta \varphi_0 = \delta \varphi_0(\mathbf{r})$$

$$=\frac{2\hbar^2}{2m}\int d^3r \,\delta\varphi_0 \cdot (-\Delta\varphi_0) + 2\int d^3r \,\delta\varphi_0 \cdot (V(r) - \mu)\varphi_0 + \frac{4}{2}(N-1)g\int d^3r \,\delta\varphi_0 \cdot \varphi_0^3$$

This results into the GP equation derived here in the variational way:

$$\frac{1}{2m}p^{2}+V(\boldsymbol{r})+(N-\boldsymbol{1})g\left|\varphi_{0}(\boldsymbol{r})\right|^{2}\right)\varphi_{0}(\boldsymbol{r})=\mu\varphi_{0}(\boldsymbol{r})$$

eliminates self-interaction

BACK

Variational estimate of the condensate properties

2 ANNEX Interpretation of the Lagrange multiplier μ

The idea is to identify it with the chemical potential. First, we modify the notation to express the particle number dependence

$$\mathcal{E}_{N}[\boldsymbol{\varphi}] = N\left\{ \left\langle \boldsymbol{\varphi} \middle| \frac{1}{2m} p^{2} \middle| \boldsymbol{\varphi} \right\rangle + \left\langle \boldsymbol{\varphi} \middle| \boldsymbol{V} \middle| \boldsymbol{\varphi} \right\rangle + \frac{1}{2} (N-1) g \int d^{3} \boldsymbol{r} \, \boldsymbol{\varphi}^{4} \right\}$$
$$E_{N} = \mathcal{E}_{N}[\boldsymbol{\varphi}_{0N}], \quad \left(\frac{1}{2m} p^{2} + V(\boldsymbol{r}) + (N-1) g \left| \boldsymbol{\varphi}_{0N}(\boldsymbol{r}) \right|^{2} \right) \boldsymbol{\varphi}_{0}(\boldsymbol{r}) = \mu_{N} \boldsymbol{\varphi}_{0N}(\boldsymbol{r})$$

The first result is that μ is not the average energy per particle:

$$E_{N} / N = \mathcal{E}_{N} [\varphi_{0N}] / N = \langle \varphi_{0N} | \frac{1}{2m} p^{2} | \varphi_{0N} \rangle + \langle \varphi_{0N} | V | \varphi_{0N} \rangle + \frac{1}{2} (N-1) g \int d^{3}r \, \varphi_{0N}^{4}$$

from the GPE $\mu_{N} = \langle \varphi_{0N} | \frac{1}{2m} p^{2} | \varphi_{0N} \rangle + \langle \varphi_{0N} | V | \varphi_{0N} \rangle + (N-1) g \int d^{3}r \, \varphi_{0N}^{4}$

Variational estimate of the condensate properties

Compare now systems with *N* and *N* -1 particles:

 $E_N = \underbrace{\mathcal{E}_N\left[\varphi_{0N}\right]}_{N-1} = \underbrace{\mathcal{E}_{N-1}\left[\varphi_{0N}\right]}_{N-1} + \mu_N \ge \underbrace{\mathcal{E}_{N-1}\left[\varphi_{0,N-1}\right]}_{N-1} + \mu_N = E_{N-1} + \mu_N$

 $\mu_N \dots$ energy to remove a particle use of the variational without relaxation of the condensate principle for GPE

In the "thermodynamic" asymptotics of large N, the inequality tends to equality. This only makes sense, and can be proved, for g > 0.

BACK

4

Reminescent of the Derivation:

theorem in the HF theory of atoms.

 $\mathcal{E}_{N}[\varphi] = N\langle \varphi | \frac{1}{2m} p^{2} | \varphi \rangle + N\langle \varphi | V | \varphi \rangle + \frac{1}{2} N(N-1)g \int d^{3}r \varphi^{4}$ $\mathcal{E}_{N-1}[\varphi] = (N-1)\langle \varphi | \frac{1}{2m} p^{2} | \varphi \rangle + (N-1)\langle \varphi | V | \varphi \rangle + \frac{1}{2} (N-1)(N-2)g \int d^{3}r \varphi^{4}$ $\mathcal{E}_{N} - \mathcal{E}_{N-1} = \underbrace{\langle \varphi | \frac{1}{2m} p^{2} | \varphi \rangle + \langle \varphi | V | \varphi \rangle + \frac{1}{2} (N(N-1) - (N-1)(N-2))g \int d^{3}r \varphi^{4}}_{\mu_{N}} \text{ for } \varphi \mapsto \varphi_{0N}$

Variational estimate of the condensate properties

③ SCALING ANSATZ FOR A SPHERICAL PARABOLIC TRAP

The potential energy has the form

$$V(\mathbf{r}) = \frac{1}{2}m\omega_0^2 \cdot r^2 = \frac{1}{2}m\omega_0^2(x^2 + y^2 + z^2)$$

Without interactions, the GPE reduces to the SE for isotropic oscillator

$$\left(\frac{1}{2m}p^2 + \frac{1}{2}m\omega_0^2 \cdot r^2\right)\varphi_0(\mathbf{r}) = \frac{3}{2}\hbar\omega_0\varphi_0(\mathbf{r})$$

The solution (for the ground state orbital) is

$$\varphi_{00}(\mathbf{r}) = A_0^3 e^{-\frac{1}{2} \cdot \frac{\mathbf{r}^2}{a_0^2}}, \quad a_0 = \sqrt{\frac{\hbar}{m\omega_0}}, \quad \hbar\omega_0 = \frac{\hbar^2}{ma_0^2} \quad A_0 = \left(a_0^2 \pi\right)^{-1/4}$$

We (have used and) will need two integrals:

$$I_{1}(\sigma) = \int_{-\infty}^{+\infty} du e^{-\frac{u^{2}}{\sigma^{2}}} = \sigma \sqrt{\pi}, \quad I_{2}(\sigma) = \int_{-\infty}^{+\infty} du e^{-\frac{u^{2}}{\sigma^{2}}} u^{2} = \frac{1}{2}\sigma^{3}\sqrt{\pi}$$

35

Variational estimate of the condensate properties

SCALING ANSATZ

The condensate orbital will be taken in the form

$$\varphi_0(\mathbf{r}) = A^3 e^{-\frac{1}{2} \cdot \frac{r^2}{b^2}}, \qquad A = (b^2 \pi)^{-1/4}$$

It is just like the ground state orbital for the isotropic oscillator, but with a rescaled size. This is reminescent of the well-known scaling for the ground state of the helium atom.

Next, the total energy is calculated for this orbital

$$\mathcal{E}[\varphi_{0}] = \frac{\hbar^{2}}{2m} N \int d^{3}r \left(\nabla \varphi_{0}(r) \right)^{2} + N \int d^{3}r V(r) \left(\varphi_{0}(r) \right)^{2} + \frac{1}{2} N \left(N - 1 \right) g \int d^{3}r \left(\varphi_{0}(r) \right)^{4}$$
$$= \frac{1}{2} \hbar \omega_{0} N A^{6} \left\{ \frac{a_{0}^{2}}{b^{4}} \int d^{3}r e^{-\frac{r^{2}}{b^{2}}} r^{2} + \frac{1}{a_{0}^{2}} \int d^{3}r e^{-\frac{r^{2}}{b^{2}}} r^{2} + (N - 1) A^{6} \frac{m a_{0}^{2}}{\hbar^{2}} g \int d^{3}r e^{-\frac{2r^{2}}{b^{2}}} \right\}$$

ADDITIONAL NOTES Variational estimate of the condensate properties For an explicit evaluation, we (have used and) will use the identities: $\frac{\hbar^2}{m} = \hbar \omega_0 a_0^2, \ m \omega_0^2 = \frac{\hbar \omega_0}{a_0^2}, \ A^2 = \frac{1}{I_1(b)} = \frac{1}{b\sqrt{\pi}}, \ g = \frac{4\pi \hbar^2 a_s}{m}$ $\mathcal{E}[\varphi_0] =$ The integrals, by the Fubini theorem, are a product of three: $=\hbar\omega_{0}N\left\{\frac{3I_{2}(b)\left(I_{1}(b)\right)^{2}}{2b\sqrt{\pi}\left(I_{1}(b)\right)^{2}}\left\{\frac{a_{0}^{2}}{b^{4}}+\frac{1}{a_{0}^{2}}\right\}+(N-1)\frac{1}{2b^{3}\pi^{3/2}}\frac{ma_{0}^{2}}{\hbar^{2}}\frac{4\pi\hbar^{2}a_{s}}{m}\frac{\left(I_{1}(b/\sqrt{2})\right)^{3}}{\left(I_{1}(b)\right)^{3}}\right\}\right\}$ Finally, $\mathcal{E}[\varphi_0] = \hbar \omega_0 N \left\{ \frac{3}{4} \right\} \left\{ \frac{a_0^2}{b^2} + \frac{b^2}{a_0^2} \right\} \left\{ + \frac{(N-1)}{\sqrt{2\pi}} \frac{a_s}{a_0} \cdot \frac{a_0^3}{b^3} \right\} \equiv \hbar \omega_0 N \cdot \tilde{E}(\tilde{\sigma})$ dimension-less energy per particle $\tilde{E}(\tilde{\sigma}) = \frac{3}{4} \left\{ \frac{1}{\tilde{\sigma}^2} + \tilde{\sigma}^2 \right\} + \eta \cdot \frac{1}{\tilde{\sigma}^3}$ $\tilde{\sigma} = \frac{b}{a}$ dimension-less orbital size

This expression is plotted in the figures in the main lecture.



37