Hartree-Fock Approximate Molecular Orbital Theory

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ture on programming

Purpose

First, we define the problem, beginning with the Schrödinger equation

$$
\ddot{H}\Psi = E\Psi. \tag{1}
$$

Our goal is to come up with an analytic equation for the energy which can be minimized with respect to some variational parameter(s) to give an upper bound on the energy. To do this, we must

- 1. Understand the hamiltonian operater H.
- 2. Find an appropriate wave function Ψ which allows simple calculation of the electronic energy.
- 3. Examine potential variational parameters to figure out how to minimize the energy.

It will likely be onvenient to have an outline of where we're going, so here it is:

- I. Puropse (whi
h we've been through)
- II. The Hamiltonian, \hat{H}
	- energy operators and
	- Coulombi potential operators

III. The Wave Function, Ψ

-
- Hartree produ
ts
- Slater determinants

IV. Hamiltonian as Energy Operator, Schrödinger Equation

- \bullet $E = \frac{1}{2}$
- Integrals over one ele
tron operators
- Integrals over two ele
tron operators
- se province experience and the expressions are associated and the set of the set of the set of the set of the s
- G corresponding to the HF energy of H

V. What Variational Parameter?

- LCAO{MO theory, energy in AO basis
- Density Matri
es

VI. Hartree-Fock Equations

- Lagrangs's Undetermined Multipliers
- A load 'o math

VII. Matrix Formalism

VIII. Program Outline

The Hamiltonian, H

The Hamiltonian is the total energy operator for a system, and is written as the sum of the kineti energy of all the omponents of the system and the internal potential energy. In an atom or molecule, comprised of positive nuclei and negative electrons, the potential energy is simply that due to the coulombic interactions present. Thus for the kinetic energy in a system of M nuclei and N electrons:

$$
\hat{T}_N = -\sum_{A}^{M} \frac{1}{2M_A} \nabla_A^2 \tag{2}
$$

$$
\hat{T}_e = -\sum_{i=1}^{N} \frac{1}{2} \nabla_i^2.
$$
\n(3)

And for the potential energy:

$$
\hat{V}_{NN} = \sum_{A>B}^{M} \frac{Z_A Z_B}{r_{AB}} \tag{4}
$$

$$
\hat{V}_{ee} = \sum_{i>j}^{N} \frac{1}{r_{ij}} \tag{5}
$$

$$
\hat{V}_{eN} = -\sum_{A}^{M} \sum_{i}^{N} \frac{Z_A}{r_{iA}}.
$$
\n(6)

 Mmnc $\mathsf{H} - \mathsf{T}$ + V,

$$
\hat{H} = -\sum_{i}^{N} \frac{1}{2} \nabla_{i}^{2} - \sum_{A}^{M} \sum_{i}^{N} \frac{Z_{A}}{r_{iA}} + \sum_{i>j}^{N} \frac{1}{r_{ij}} - \sum_{A}^{M} \frac{1}{2M_{A}} \nabla_{A}^{2} + \sum_{A>B}^{M} \frac{Z_{A}Z_{B}}{r_{AB}}.
$$
\n(7)

Within the Born-Oppenheimer approximation, we assume the nuclei are held fixed while the electrons move really fast around them. (note: $M_p/M_e \approx 1840$.) In this case, nuclear motion and ele
troni motion are seperated. The last two terms an be removed from the total hamiltonian to give the electronic namiltonian, H_e , since $v_{NN} = K$, and $v_A^2 = 0$. The nuclear motion is handled in a rotational/vibrational analysis. We will be working within the $B-O$ approximation, so realizing

$$
\hat{H}_e = -\sum_{i}^{N} \frac{1}{2} \nabla_i^2 - \sum_{A}^{M} \sum_{i}^{N} \frac{Z_A}{r_{iA}} + \sum_{i>j}^{N} \frac{1}{r_{ij}}
$$
\n(8)

we completly define the problem. Solving the electronic Schrödinger equation using this will give the electronic structure of a molecular system at a fixed nuclear geometry.

The Wave Function, Ψ

We've derived a complete many-electron Hamiltonian operator. Of course, the Schrödinger equation involving it is intractable, so let's consider a simpler problem, involving the one-electron hamiltonian

$$
\hat{h}(i) = -\frac{1}{2}\nabla_i^2 + \sum_{A}^{M} \frac{1}{r_{iA}}\tag{9}
$$

which involves no electron-electron interaction. This is soluable in the B-O approximation (recall the hydrogen atom by letting $M = 1$). Call the solutions to the one-electron Schrödinger equation $\chi_i(\mathbf{x})$. These will be molecular spin orbitals when we get around to it, but for now let it suffice to know they satisify the eigenequation

$$
h(i)\chi_j(\mathbf{x}_i) = \varepsilon_j \chi_j(\mathbf{x}_i) \tag{10}
$$

with the interpretation that electron i occupies spin orbital χ_i with energy ε_i . If we ignore electron—electron interaction in n_e , we construct a simpler system with Hamiltonian

$$
\hat{H} = \sum_{i}^{N} \hat{h}(i). \tag{11}
$$

It will have eigenfunctions which are simple products of occupied spin orbitals, and thus an energy which is a sum of individual orbital energies, as

$$
\Psi^{HP} = \chi_i(\mathbf{x}_1)\chi_j(\mathbf{x}_2)\chi_k(\mathbf{x}_3)\cdots\chi_{\eta}(\mathbf{x}_N) \tag{12}
$$

$$
\langle \Psi | \hat{H} | \Psi \rangle = \varepsilon_i + \varepsilon_j + \varepsilon_k + \ldots + \varepsilon_\eta = E. \tag{13}
$$

This kind of wavefunction is called a Hartree Product, and it is not physically realistic. In the first place, it is an independent-electron model, and we know electrons repel each other. Secondly, it does not satisfy the antisymmetry principle due to Pauli which states that the sign of the wavefunction must be inverted under the operation of switching the coordinates of any two ele
trons, or

$$
\Psi(\cdots \mathbf{x}_i \cdots \mathbf{x}_j \cdots) = -\Psi(\cdots \mathbf{x}_j \cdots \mathbf{x}_i \cdots).
$$
\n(14)

Part of the proof of equation 13 acknowledges this is not so for a Hartree Product. To remedy this, first consider a two-electron system, such as helium. Two equivalent Hartree Product wavefun
tions for this system are

$$
\Psi_1^{HP} = \chi_i(\mathbf{x}_1)\chi_j(\mathbf{x}_2) \n\Psi_2^{HP} = \chi_i(\mathbf{x}_2)\chi_j(\mathbf{x}_1).
$$
\n(15)

Obviously, neither of these is appropriate. However, using the old "by inspection..." trick, we noti
e that

$$
\Psi = \frac{1}{\sqrt{2}} [\chi_i(\mathbf{x}_1)\chi_j(\mathbf{x}_2) - \chi_i(\mathbf{x}_2)\chi_j(\mathbf{x}_1)] \tag{16}
$$

does. The mathematical form of this wavefunction can be generated by a determinant of χ 's,

$$
\Psi = 2^{-1/2} \begin{vmatrix} \chi_i(\mathbf{x}_1) & \chi_j(\mathbf{x}_1) \\ \chi_i(\mathbf{x}_2) & \chi_j(\mathbf{x}_2) \end{vmatrix} . \tag{17}
$$

The familiar Pauli exclusion principle follows directly from this example. When we attempt to doubly occupy a spin orbital χ_i by putting electron 1 and electron 2 in it, what happens?

$$
\Psi = 2^{-1/2} \begin{vmatrix} \chi_i(\mathbf{x}_1) & \chi_i(\mathbf{x}_1) \\ \chi_i(\mathbf{x}_2) & \chi_i(\mathbf{x}_2) \end{vmatrix}
$$
\n
$$
= \frac{1}{\sqrt{2}} [\chi_i(\mathbf{x}_1) \chi_i(\mathbf{x}_2) - \chi_i(\mathbf{x}_2) \chi_i(\mathbf{x}_1)]
$$
\n
$$
= 0
$$
\n(18)

Equation 17 can be generalized to give the N electron Slater determinant

$$
\Psi = (N!)^{-1/2} \begin{vmatrix} \chi_i(\mathbf{x}_1) & \chi_j(\mathbf{x}_1) & \cdots & \chi_{\eta}(\mathbf{x}_1) \\ \chi_i(\mathbf{x}_2) & \chi_j(\mathbf{x}_2) & \cdots & \chi_{\eta}(\mathbf{x}_2) \\ \vdots & \vdots & \ddots & \vdots \\ \chi_i(\mathbf{x}_N) & \chi_j(\mathbf{x}_N) & \cdots & \chi_{\eta}(\mathbf{x}_N) \end{vmatrix} .
$$
\n(19)

A shorthand notation for a Slater determinant has been introdu
ed, where all the diagonal elements in the determinant are written in order as a "ket" vector. Equation 19 can thus be written as

$$
\Psi = |\chi_i(\mathbf{x}_1)\chi_j(\mathbf{x}_2)\cdots\chi_\eta(\mathbf{x}_N)\rangle \tag{20}
$$

where the normalization constant is absorbed into the notation.

Now we have written down a wave function appropriate for use in the case where $H = \sum h(i)$. In HF theory, we make some simplifications so many-electron atoms and molecules can be treated this way. By tacitly assuming that each electron moves in a percieved electric field generated by the stationary nuclei and the average spatial distribution of all the other electrons, it essentially becomes an independant-electron problem. The HF Self Consistent Field procedure (SCF) will be bent on constructing each $\chi(\mathbf{x})$ to give the lowest energy.

Energy Expressions

Let's assume a wave function of the Slater determinant form and find an expression for the expectation value of the energy. We've written a Slater determinant as a ket vector in shorthand notation, allowing us to make use of Dirac notation for such things as overlap. In this context, re
all that

$$
\langle \Psi_a | \Psi_b \rangle = (\Psi_a, \Psi_b) \n\langle \Psi_a | \Psi_b \rangle = \int \Psi_a^*(\mathbf{x}) \Psi_b(\mathbf{x}) d\mathbf{x}
$$
\n(21)

where the basis vectors Ψ is expanded in are every possible value of x with contraction coefficients α in a contract the value of \bm{x} of \bm{x} . Thus placing an operator (such as H) inside the bracket, we α and expectation value of the observable associated with that operator. Since H is the energy operator,

$$
E_0 = \int d\tau \Psi_0^* \hat{H}_e \Psi_0
$$

= $\langle \Psi_0 | \hat{H}_e | \Psi_0 \rangle.$ (22)

 $d\tau$ is the differential of all the spin and space coordinates of all the electrons.

with much foresignt, we continue to simplify the problem by writing H as a sum of one and two-electron operators

$$
\hat{H}_e = \sum_i \hat{h}(i) + \sum_{i>j} \frac{1}{r_{ij}} \tag{23}
$$

$$
= \hat{H}_1^{core} + \hat{H}_2. \tag{24}
$$

This will allow us to more precisely develop the electronic energy by it's components.

First, examine the core hamiltonian H_1^{core} .

$$
\langle \Psi | \hat{H}_1^{core} | \Psi \rangle = \sum_i \langle \Psi | \hat{h}(i) | \Psi \rangle \tag{25}
$$

$$
= \sum_{i} \langle \chi_1(\mathbf{x}_1) \dots \chi_i(\mathbf{x}_i) \dots | \hat{h}(i) | \chi_1(\mathbf{x}_1) \dots \chi_i(\mathbf{x}_i) \dots \rangle \qquad (26)
$$

The nature of this is best evidenced by example, so we turn to the familiar helium atom, $\Psi =$ $\frac{|\chi_1(\mathbf{A})|}{\chi_2(\mathbf{A}_2)}$. Look at one term in the above sum, for the sake of inustration take $\frac{h(1)}{h(1)}$.

$$
\langle \chi_1(\mathbf{x}_1)\chi_2(\mathbf{x}_2)|\hat{h}(1)|\chi_1(\mathbf{x}_1)\chi_2(\mathbf{x}_2)\rangle = \frac{1}{2}\int [\chi_1^*(\mathbf{x}_1)\chi_2^*(\mathbf{x}_2) - \chi_1^*(\mathbf{x}_2)\chi_2^*(\mathbf{x}_1)]\hat{h}(1) [\chi_1(\mathbf{x}_1)\chi_2(\mathbf{x}_2) - \chi_1(\mathbf{x}_2)\chi_2(\mathbf{x}_1)]d\mathbf{x}_1d\mathbf{x}_2
$$
\n(27)

$$
= \frac{1}{2} \int \chi_1^*(\mathbf{x}_1) \chi_2^*(\mathbf{x}_2) \hat{h}(1) \chi_1(\mathbf{x}_1) \chi_2(\mathbf{x}_2) d\tau + \int \chi_1^*(\mathbf{x}_2) \chi_2^*(\mathbf{x}_1) \hat{h}(1) \chi_1(\mathbf{x}_2) \chi_2(\mathbf{x}_1) d\tau - \int \chi_1^*(\mathbf{x}_2) \chi_2^*(\mathbf{x}_1) \hat{h}(1) \chi_1(\mathbf{x}_1) \chi_2(\mathbf{x}_2) d\tau - \int \chi_1^*(\mathbf{x}_1) \chi_2^*(\mathbf{x}_2) \hat{h}(1) \chi_1(\mathbf{x}_2) \chi_2(\mathbf{x}_1) d\tau
$$
(28)

$$
- \int \chi_1(\mathbf{x}_2) \chi_2(\mathbf{x}_1) \mu(1) \chi_1(\mathbf{x}_1) \chi_2(\mathbf{x}_2) u_1 - \int \chi_1(\mathbf{x}_1) \chi_2(\mathbf{x}_2) \mu(1) \chi_1(\mathbf{x}_2) \chi_2(\mathbf{x}_1) u_1 \qquad (26)
$$

= $\frac{1}{2} (\langle \chi_1 | \hat{h} | \chi_1 \rangle + \langle \chi_2 | \hat{h} | \chi_2 \rangle)$ (29)

Here $\langle \chi_1|h|\chi_1\rangle$ is defined as $\int \chi_1^*(\mathbf{x}_1)h(1)\chi_1(\mathbf{x}_1)d\mathbf{x}_1$. In the first two terms of equation 28, the integrations over electron two's coordinates can be carried out irrespective of electron one's, and give the two terms of equation 29. The last two terms integrate to zero due to the orthogonality of χ_1 and χ_2 . Repeating this for $n(z)$ we get exactly the same thing, and we see

$$
\sum_{i} \langle \Psi | \hat{h}(i) | \Psi \rangle = \langle \chi_1 | \hat{h} | \chi_1 \rangle + \langle \chi_2 | \hat{h} | \chi_2 \rangle \tag{30}
$$

Profound, isn't it? Seems that every occupied spin orbital χ_i yields a term of the form $\langle \chi_i|h \chi_i \rangle$ to the one electron energy.

 1 vow loon at π_2 .

$$
\langle \Psi | \hat{H}_2 | \Psi \rangle = \sum_{i > j} \langle \Psi | \frac{1}{r_{ij}} | \Psi \rangle
$$

=
$$
\sum_{i > j} \langle \chi_1(\mathbf{x}_1) \dots \chi_i(\mathbf{x}_i) \dots | \frac{1}{r_{ij}} | \chi_1(\mathbf{x}_1) \dots \chi_i(\mathbf{x}_i) \dots \rangle
$$
(31)

Continuing to work in the helium atom example (realize that this could be any two electron system) pick $(i, j) = (1, 2)$ and look at that one term.

$$
\langle \chi_1(\mathbf{x}_1)\chi_2(\mathbf{x}_2)|\frac{1}{r_{12}}|\chi_1(\mathbf{x}_1)\chi_2(\mathbf{x}_2)\rangle = \frac{1}{2}\int [\chi_1(\mathbf{x}_1)\chi_2(\mathbf{x}_2) - \chi_1(\mathbf{x}_2)\chi_2(\mathbf{x}_1)]\frac{1}{r_{12}}\n\quad [\chi_1(\mathbf{x}_1)\chi_2(\mathbf{x}_2) - \chi_1(\mathbf{x}_2)\chi_2(\mathbf{x}_1)]d\mathbf{x}_1d\mathbf{x}_2
$$
\n(32)

$$
= \frac{1}{2}\int \chi_1^*(\mathbf{x}_1)\chi_2^*(\mathbf{x}_2)\frac{1}{r_{12}}\chi_1(\mathbf{x}_1)\chi_2(\mathbf{x}_2)d\tau + \int \chi_1^*(\mathbf{x}_2)\chi_2^*(\mathbf{x}_1)\frac{1}{r_{12}}\chi_1(\mathbf{x}_2)\chi_2(\mathbf{x}_1)d\tau -\int \chi_1^*(\mathbf{x}_2)\chi_2^*(\mathbf{x}_1)\frac{1}{r_{12}}\chi_1(\mathbf{x}_1)\chi_2(\mathbf{x}_2)d\tau - \int \chi_1^*(\mathbf{x}_1)\chi_2^*(\mathbf{x}_2)\frac{1}{r_{12}}\chi_1(\mathbf{x}_2)\chi_2(\mathbf{x}_1)d\tau
$$
(33)

Unfortunately, the $\frac{1}{r_{12}}$ operator prevents seperation of the integrations over the electronic coor-

general, they are not. However, since the x_1 and x_2 are dummy variables, the first and second terms of equation 33 are equal, as are the last two. Thus for the two electron operator $\frac{1}{r_{12}},$

$$
\langle \chi_1(\mathbf{x}_1)\chi_2(\mathbf{x}_2)|\frac{1}{r_{12}}|\chi_1(\mathbf{x}_1)\chi_2(\mathbf{x}_2)\rangle = \langle 12|12\rangle - \langle 12|21\rangle = \langle 12||12\rangle \tag{34}
$$

where

$$
\langle ij|kl\rangle = \int d\mathbf{x}_1 d\mathbf{x}_2 \chi_i^*(\mathbf{x}_1) \chi_j^*(\mathbf{x}_2) \frac{1}{r_{12}} \chi_k(\mathbf{x}_1) \chi_l(\mathbf{x}_2)
$$
(35)

$$
\langle ij||kl\rangle = \langle ij|kl\rangle - \langle ij|lk\rangle. \tag{36}
$$

The construction $\langle ij||kl\rangle$ is called an antisymmetrized two electron integral in physicists notation.

By working in the spin orbital basis, much trouble is avoided. In fact, by extending the results shown previously to the general case, we can now write down the HF energy for a given set of o

upied spin orbitals.

$$
E_{HF} = \langle \Psi | \hat{H}_e | \Psi \rangle = \langle \Psi | \hat{H}_1^{core} + \hat{H}_2 | \Psi \rangle \tag{37}
$$

$$
E_{HF} = \langle \chi_1(\mathbf{x}_1) \dots \chi_i(\mathbf{x}_i) \dots | \hat{H}_1 | \chi_1(\mathbf{x}_1) \dots \chi_i(\mathbf{x}_i) \dots \rangle + \n\langle \chi_1(\mathbf{x}_1) \dots \chi_i(\mathbf{x}_i) \dots | \hat{H}_2 | \chi_1(\mathbf{x}_1) \dots \chi_i(\mathbf{x}_i) \dots \rangle = \sum \langle \chi_1(\mathbf{x}_1) \dots \chi_i(\mathbf{x}_i) \dots | \hat{h}(i) | \chi_1(\mathbf{x}_1) \dots \chi_i(\mathbf{x}_i) \dots \rangle +
$$
\n(38)

$$
\sum_{i>j}^{i} \langle \chi_1(\mathbf{x}_1) \dots \chi_i(\mathbf{x}_i) \dots | \frac{1}{r_{12}} | \chi_1(\mathbf{x}_1) \dots \chi_i(\mathbf{x}_i) \dots \rangle \tag{39}
$$

$$
= \sum_{i} \langle \chi_i | h | \chi_i \rangle + \sum_{i > j} \langle i j | | i j \rangle \tag{40}
$$

Now move on and consider working in the spatial orbital basis, where

$$
\chi_i(\mathbf{x}) = \psi_a(\mathbf{r})\omega. \tag{41}
$$

This is more natural, since our intuition is usually based on having a region of space which describes the location (more or less) of <u>two</u> electrons, one of alpha spin and one of beta spin. Some of quantum hemistry is formulated entirely in terms of spin orbitals, for various reasons. For our purposes, we will work entirely in the spatial orbital basis. This will cause things to get somewhat murky soon, but in the long run it will be simpler.

At any rate, in the two electron system we adore so much, we can identify the two occupied spin orbitals χ_1 and χ_2 as the spin up and spin down halves of the single lowest lying spatial orbital, a 1s in helium or the σ bonding orbital in H_2 for example. These can be more precisely defined as

$$
\chi_1(\mathbf{x}) = \psi_1(\mathbf{r}) \alpha
$$

$$
\chi_2(\mathbf{x}) = \psi_1(\mathbf{r})\beta. \tag{42}
$$

(43)

This changes the way we write slater determinants. Using an overbar to denote β spin occupation of a spatial orbital, $\overline{\psi}$,

$$
|\chi_1(\mathbf{x}_1)\chi_2(\mathbf{x}_2)\rangle = |\psi_1(\mathbf{r}_1)\overline{\psi}_1(\mathbf{r}_2)\rangle = |\psi_1\overline{\psi}_1\rangle
$$
\n(44)

Rethinking the one electron integrals for this case,

$$
\sum_{i} \langle \psi_{1} \overline{\psi}_{1} | \hat{h}(i) | \psi_{1} \overline{\psi}_{1} \rangle = \langle \psi_{1} | \hat{h}(1) | \psi_{1} \rangle + \langle \overline{\psi}_{1} | \hat{h}(2) | \overline{\psi}_{1} \rangle \tag{45}
$$

 $-$ (ψ_1 | ψ_1 | ψ_1 | ψ_1 | ψ_1 j 1) (and $\sqrt{-1}$)

$$
\equiv h_{11} + h_{11} \tag{47}
$$

The notation $(\psi_1|\mu|\psi_1)$ is used to denote an integral over only spatial coordinates, what remains after the spin integrations have been carried out, giving a factor of 1 or 0.

That was a neat closed shell system. How about something like $\Psi = |\psi_1 \overline{\psi}_1 \psi_2\rangle$?

$$
\sum_{i} \langle \Psi | \hat{h}(i) | \Psi \rangle = \langle \psi_1 | \hat{h} | \psi_1 \rangle + \langle \overline{\psi}_1 | \hat{h} | \overline{\psi}_1 \rangle + \langle \psi_2 | \hat{h} | \psi_2 \rangle \tag{48}
$$

$$
= (\psi_1|\hat{h}|\psi_1) + (\psi_1|\hat{h}|\psi_1) + (\psi_2|\hat{h}|\psi_2)
$$

= 2h₁₁ + h₂₂ (49)

$$
= 2\sum_{i}^{docc} h_{ii} + \sum_{i}^{socc} h_{ii}
$$

$$
= 2\sum_{i}^{occ} f_{i}h_{ii}
$$
 (50)

The coefficient f_i here is related to the occupation of spatial orbital i, and will be more precisely defined later. The two electron integrals are a tad more involved, but we go about it in essentially the same manner.

$$
\langle \chi_1 \chi_2 | \chi_1 \chi_2 \rangle = \langle \psi_1 \overline{\psi}_1 | \psi_1 \overline{\psi}_1 \rangle \tag{51}
$$

$$
= \left[\psi_1 \psi_1 | \psi_1 \psi_1 \right] \tag{52}
$$

$$
= \left(\psi_1 \psi_1 | \psi_1 \psi_1 \right) \tag{53}
$$

$$
\langle \chi_1 \chi_2 | \chi_2 \chi_1 \rangle = \langle \psi_1 \psi_1 | \psi_1 \psi_1 \rangle
$$
\n
$$
= [\psi_1 \overline{\psi}, \overline{\psi}, \psi_1] \tag{55}
$$

$$
= \left[\psi_1 \psi_1 | \psi_1 \psi_1 \right] \tag{55}
$$

$$
= 0 \tag{56}
$$

Yes, I know. Very confusing. But it's all just notation, and can be understood. In physicist's notation (equivalent to Dirac notation), $\langle \psi_i \psi_j | \psi_k \psi_l \rangle$ refers to the two electron integral where ψ_i and ψ_k are functions of electron 1, while ψ_j and ψ_l are functions of electron 2. Chemist's notation (with the square brackets []) places the functions of electron 1 on the left and the functions of electron 2 on the right. When the two functions of a single electron are not of the same spin, the whole integral goes to zero, otherwise the spin integrates out to 1. Hence the spin-free notation $(\psi_1 \psi_1 | \psi_1 \psi_1)$. What occurs when the two electrons are of parallel spin, requiring distinct spatial orbitals and a wavefunction something like $\Psi = |\psi_1 \psi_2\rangle$? The same general form is present, and a related antisymmetrized two electron integral is evaluated. In this case, $\langle \psi_1 \psi_2 | \psi_1 \psi_2 \rangle$.

$$
\langle \psi_1 \psi_2 | | \psi_1 \psi_2 \rangle = \langle 12 | | 12 \rangle
$$

= $\langle 12 | 12 \rangle - \langle 12 | 21 \rangle$
= $[11 | 22] - [12 | 21]$
= $(11 | 22) - (12 | 21)$ (57)

$$
\langle \psi_1 \psi_2 | \psi_1 \psi_2 \rangle = J_{12} - K_{12} \tag{58}
$$

Another bit of notation, which should be apparent. $J_{12} = (11|22)$ and $K_{12} = (12|21)$. J_{ij} is termed a coulomb integral and has the physically reassuring interpretation of somehow accounting for electronic repulsion between electrons in molecular orbital i and molecular orbital j. K_{ij} , the exchange integral, has no classical analog and no true physical interpretation. Many have tried to come up with something, and a typical attempt says that it "correlates the motions of electron i and j when they have parallel spins, lowering the energy since those electrons avoid each other better." Whatever.

Best to just move on to a general energy equation in the spatial MO basis. In summary:

- \bullet The one electron integrals contribute n_{ii} for each electron in orbital $i.$
- The two electron integrals contribute J_{ij} for each pair of electrons, and $(-K_{ij})$ for each pair of parallel spin.

$$
E_{HF} = 2 \sum_{i}^{occ} f_{i} h_{ii} + \sum_{i,j}^{occ} {\{\alpha_{ij}(ii|jj) + \beta_{ij}(ij|ij)\}}
$$

=
$$
2 \sum_{i}^{occ} f_{i} h_{ii} + \sum_{i,j}^{occ} {\{\alpha_{ij} J_{ij} + \beta_{ij} K_{ij}\}}
$$
(59)

$$
f_i = \begin{cases} 1 \text{ if } i \text{ doubly occupied} \\ 1/2 \text{ if } i \text{ singly occupied} \end{cases}
$$
\n
$$
\alpha_{ij} = \begin{cases} 2 \text{ if } i \text{ and } j \text{ doubly occupied} \\ 1 \text{ if } i \text{ or } j \text{ doubly occupied and the other singly occupied} \\ 1/2 \text{ if } i \text{ and } j \text{ singly occupied} \end{cases}
$$
\n(61)

 μ_{ij} $-$ ⁸ \perp \perp -1 if it is in the contract of upied, the other double in the other document of the other double the other double in the other double that th -1/2 if ⁱ and ^j singly o

upied, with parallel spin 1/2 if ⁱ and ^j singly o

upied, with opposite spin

(62)

What Variational Parameter?

Ah, the crux of the problem, is it not? Up until now, we've just assumed we have some set of molecular orbitals χ_i or ψ_i , which we can manipulate at will. But how does one come up with even approximate solutions to the many body Schrödinger equation without having to solve it? Start with the celebrated linear combination of atomic orbitals to get molecular orbitals (LCAO– MO) approximation. This allows us to use some set of (approximate) atomi orbitals, the basis functions which we know and love, to expand the MOs in. In the most general terms,

$$
\psi_i = \sum_{\mu} C^i_{\mu} \phi_{\mu}.
$$
\n(63)

 ψ_i remains a spatial molecular orbital, ϕ_μ is a spatial atomic orbital (perhaps symmetry orbital, but no matter), and C^i_μ are the contraction coefficients by which we transform from one basis to another. Armed with only this, we should be able to compose the electronic energy in the atomic orbital basis. Why, you ask? Be
ause we have an expression in terms of integrals over MOs. To variationally minimize that energy, we need to vary the MOs themselves, but have no way to do that, since they remain these amorphous constructs. By defining them a bit more precisely, we should arrive at a point where an obvious set of variational parameters (hint: C^i_μ) present themselves. Begin with the closed shell HF energy in terms of spatial MOs.

$$
E_{HF} = 2 \sum_{i}^{docc} h_{ii} + \sum_{i,j}^{docc} \{2J_{ij} - K_{ij}\}\
$$

=
$$
2 \sum_{i}^{docc} h_{ii} + \sum_{i,j}^{docc} \{2(ii|jj) - (ij|ij)\}\
$$

Since $(\psi_i|h|\psi_i) = \int \psi_i^*(\mathbf{r}_1)h(1)\psi_i(\mathbf{r}_1)d\mathbf{r}_1$ and $\psi_i = \sum_{\mu} C^i_{\mu} \phi_{\mu}$,

$$
\int \psi_i^*(\mathbf{r}_1)h(1)\psi_i(\mathbf{r}_1)d\mathbf{r}_1 = \int \sum_{\mu} C_{\mu}^{i*} \phi_{\mu}^*(\mathbf{r}_1)h(1) \sum_{\nu} C_{\nu}^{i} \phi_{\nu}(\mathbf{r}_1)d\mathbf{r}_1
$$
\n
$$
= \sum_{\mu} \sum_{\nu} C_{\mu}^{i*} C_{\nu}^{i} \int \phi_{\mu}^*(\mathbf{r}_1)h(1)\phi_{\nu}(\mathbf{r}_1)d\mathbf{r}_1
$$
\n
$$
= \sum_{\mu} \sum_{\nu} C_{\mu}^{i*} C_{\nu}^{i}(\phi_{\mu}|h|\phi_{\nu})
$$
\n(64)

$$
= \sum_{\mu} \sum_{\nu} C_{\mu}^{i*} C_{\nu}^{i} h_{\mu\nu} \tag{65}
$$

 $(\phi_{\mu} | h | \phi_{\nu})$ is a one electron integral over atomic orbitals. Do we have something we can actually calculate!? Take an aside and examine this quantity superficially.

Typi
ally, basis fun
tions are onstru
ted to mimi true atomi orbitals. The Hydrogen atom can be described rigorously, and the eigenfunctions found. The 1s orbital looks something like $Ne^{-\zeta r}$. It satisfies all the appropriate boundary conditions, having a cusp at the nucleus and exponentially decaying to zero at infinity. Higher angular momentum functions, like 2p's and 3d's, can be built from this basic framework through adding the angular nodes by multiplying in factors of x, y, and z. Basis functions such as these are called slater-type orbitals. If instead of the exponential $Ne^{-\zeta r}$ we use a gaussian function, $Ne^{-\alpha r^2}$, we loose the boundary conditions but generate a more tractable problem when it comes to calculating integrals. Using a linear combination of single artesian gaussian{type orbitals to approximate a slater{type orbital gives better computational accuracy without too much more effort. Here's the functional forms of all three types:

$$
\phi_{\mu}^{STO}(\mathbf{r}) = Nx^l y^m z^n e^{-\zeta r} \tag{66}
$$

$$
\phi_{\nu}^{SGTO}(\mathbf{r}) = Nx^{l}y^{m}z^{n}e^{-\alpha r^{2}}
$$
\n(67)

$$
\phi_{\mu}^{CGTO}(\mathbf{r}) = \sum_{\nu} C_{\nu}^{\mu} \phi_{\nu}^{SGTO}(\mathbf{r}) \tag{68}
$$

Just taking a 1s SGTO for illustrative purposes, what is that one ele
tron integral?

$$
(\phi_{1s}^{SGTO}(\mathbf{r}_1)|h(1)|\phi_{1s}^{SGTO}(\mathbf{r}_1)) = N^2 \int e^{-\alpha r^2} (-\nabla^2 + \frac{1}{r}) e^{-\alpha r^2} d\mathbf{r}
$$
(69)

Hey! We an do that!

Ba
k to the problem at hand, we now need to expand the two ele
tron integrals in the MO basis. Following a pro
edure analogus to equation 64, we get

$$
(ij|kl) = \sum_{\mu} \sum_{\nu} \sum_{\rho} \sum_{\sigma} C_{\mu}^{i*} C_{\nu}^{j*} C_{\rho}^{k} C_{\sigma}^{l} (\mu \nu | \rho \sigma)
$$
\n⁽⁷⁰⁾

All of these AO integrals can be calculated and stored, to be called up when needed to evaluate the ele
troni energy. The losed shell energy in the AO basis an be written as

$$
E^{\text{clscf}} = 2 \sum_{\mu\nu} \mathbf{D}_{\mu\nu} h_{\mu\nu} + \sum_{\mu\nu\rho\sigma} \mathbf{D}_{\mu\nu} \mathbf{D}_{\rho\sigma} \{ 2(\mu\nu|\rho\sigma) - (\mu\rho|\nu\sigma) \}
$$
(71)

$$
= \sum_{\mu\nu} \mathbf{D}_{\mu\nu} [2h_{\mu\nu} + \sum_{\rho\sigma} \mathbf{D}_{\rho\sigma} \{2(\mu\nu|\rho\sigma) - (\mu\rho|\nu\sigma)\}] \tag{72}
$$

 $\mathbf{D}_{\mu\nu}$ is the density matrix, a product of AO–MO coefficient matrices, or

$$
\mathbf{D}_{\mu\nu} = \sum_{\mu\nu} C_{\mu}^{i*} C_{\nu}^{i}
$$

$$
\mathbf{D}_{\mu\nu} = \mathbf{C}^{\dagger} \mathbf{C}
$$
 (73)

Hartree Fo
k Equations

The electronic energy is a functional of the spin orbitals, and we want to minimize it subject to some set of constraints. This can be done using the calculus of variations applied to functionals. So lets look at a general example of functional variation applied to E , a functional of some trial wavefunction Φ that can be linearly varied under a single constraint.

$$
E = \langle \Phi | H | \Phi \rangle \tag{74}
$$

$$
|\Phi\rangle = \sum_{i}^{N} c_{i} |\Psi_{i}\rangle \tag{75}
$$

By equation 74, we see that $E[\Phi]$, depending on the form of the wavefunction, and by equation 75 that $|\Phi\rangle$ can be linearly expanded (hence linearly varied) in some set of N functions. This is directly analogous to expanding the asymmetric top rotational wavefunctions in a complete set of symmetric top rotational wavefunctions. The task is to minimize E subject to the single constraint that the wavefunction Φ remain normalized, or

$$
\langle \Phi | \Phi \rangle = \sum_{ij} c_i^* c_j \langle \Psi_i | \Psi_j \rangle = 1 \tag{76}
$$

Writing the energy as

$$
E = \langle \Phi | H | \Phi \rangle = \sum_{ij} c_i^* c_j \langle \Psi_i | H | \Psi_j \rangle \tag{77}
$$

we want $\delta E = 0$, so

$$
\delta E = \frac{\partial E}{\partial c_1} \delta c_1 + \frac{\partial E}{\partial c_2} \delta c_2 + \frac{\partial E}{\partial c_3} \delta c_3 + \dots + \frac{\partial E}{\partial c_N} \delta c_N
$$

=
$$
\sum_{i}^{N} \frac{\partial E}{\partial c_i} \delta c_i = 0.
$$
 (78)

However, due to the normalization constraint, there is one linear dependancy in the expansion coefficients. If we simply solve equation 78 for the coefficients, they may no longer be normalized, and if we solve for $N-1$ of the coefficients and invoke the normalization condition to determine the N^{th} , the energy may not be stationary about it. So we construct the function

$$
\mathcal{L}[\Phi] = \langle \Phi | H | \Phi \rangle + E(\langle \Phi | \Phi \rangle - 1) \n\mathcal{L}[\Phi] = \sum_{ij} c_i^* c_j \langle \Psi_i | H | \Psi_j \rangle + \sum_{ij} c_i^* c_j E(\langle \Psi_i | \Psi_j \rangle - 1).
$$
\n(79)

Taking the differential of \mathcal{L}_1

$$
\delta \mathcal{L}[\Phi] = \sum_{ij} \delta c_i^* c_j \langle \Psi_i | H | \Psi_j \rangle + \sum_{ij} c_i^* \delta c_j \langle \Psi_i | H | \Psi_j \rangle +
$$

$$
E \sum_{ij} \delta c_i^* c_j \langle \Psi_i | \Psi_j \rangle + E \sum_{ij} c_i^* \delta c_j \langle \Psi_i | \Psi_j \rangle
$$

\n
$$
= \sum_{ij} \delta c_i^* c_j [\langle \Psi_i | H | \Psi_j \rangle + E \langle \Psi_i | \Psi_j \rangle]
$$

\n
$$
+ \sum_{ij} c_i^* \delta c_j [\langle \Psi_i | H | \Psi_j \rangle + E \langle \Psi_i | \Psi_j \rangle]
$$

\n
$$
= \sum_{i} \delta c_i^* [\sum_{j} c_j (\langle \Psi_i | H | \Psi_j \rangle + E \langle \Psi_i | \Psi_j \rangle)] +
$$

\n
$$
\sum_{j} \delta c_j [\sum_{i} c_i^* (\langle \Psi_i | H | \Psi_j \rangle + E \langle \Psi_i | \Psi_j \rangle)]]
$$

\n
$$
= \sum_{i} \delta c_i^* [\sum_{j} c_j (\langle \Psi_i | H | \Psi_j \rangle + E \langle \Psi_i | \Psi_j \rangle)] +
$$

\n
$$
\sum_{i} \delta c_i [\sum_{j} c_j^* (\langle \Psi_j | H | \Psi_i \rangle + E \langle \Psi_j | \Psi_i \rangle)]
$$

\n
$$
= \sum_{i} \delta c_i^* [\sum_{j} c_j (H_{ij} + ES_{ij}] + \sum_{i} \delta c_i [\sum_{j} c_j^* (H_{ij}^* + ES_{ij}^*)] = 0
$$
(81)

Since oc_i and oc_i are arbitrary, the bracketed parts of equation 81 must be zero. Thus

$$
\sum_{j} (H_{ij}c_j - ES_{ij}c_j) = 0.
$$
\n(82)

It is clear that this can be written as a matrix product, and is in fact an eigenvalue equation in the form

$$
\mathbf{Hc} = \mathbf{Sc}E\tag{83}
$$

Knowing that H and S are hermetian, this matrix eigenvalue equation can be rewritten as

$$
\mathbf{S}^{-1/2}\mathbf{Hc} = \mathbf{S}^{1/2}\mathbf{c}E
$$

$$
\mathbf{S}^{-1/2}\mathbf{H}\mathbf{S}^{-1/2}\mathbf{S}^{1/2}\mathbf{c} = \mathbf{S}^{1/2}\mathbf{c}E
$$
(84)

The matrix $S^{-1/2}HS^{-1/2}$ is symmetric and easily diagonalized, with eigenvectors $S^{1/2}c$. These can be transformed (multiply on the left by $S^{-1/2}$) to give the optimal coefficients c_i for each stationary state. This is very powerful, since in one fell swoop we've got the entire energy spectrum and the appropriate wave functions, properly orthonormal, for all the states. This should illustrate the general technique we will be employing to develop the Hartree–Fock equations and from them the algebrai Roothaan equations, whi
h you will be programming later this summer.

On to the true problem. Assume we have a wave function in the form of a Slater determinant of spin orbitals, $\Psi_0 = |\chi_1 \chi_2 \cdots \chi_N\rangle$. We state the problem as:

Please minimize the electronic energy of this single determinant subject to the constraint that the spin orbitals all remain orthonormal to one another.

We already understand the energy fine by equation 37, and the constraint can be simply stated as

$$
\langle a|b\rangle - \delta_{ab} = 0\tag{85}
$$

There are *N* spin orbitals, so there are *N(N* + 1)/2 independent constraints (note: $\langle a|o\rangle = \langle o|a\rangle$), so we need that many undetermined multipliers in our lagrangian function, for which we present

$$
\mathcal{L}[\{\chi_a\}] = E_0[\{\chi_a\}] - \sum_{a}^{N} \sum_{b}^{a} \varepsilon_{ba} (\langle a|b \rangle - \delta_{ab}) \tag{86}
$$

The restricted sum will prove inconvenient, but it can be eliminated. By taking the complex conjugate of the constraints and the lagrangian function, equation 86.

$$
\langle a|b\rangle^* - \delta_{ab} = 0\langle b|a\rangle - \delta_{ab} = 0
$$
\n(87)

$$
\mathcal{L}[\{\chi_a\}] = E_0[\{\chi_a\}] - \sum_{a}^{N} \sum_{b}^{a} \varepsilon_{ba}^*(\langle b|a\rangle - \delta_{ab}) \tag{88}
$$

we realize we can unrestrict the summation by restraining ε_{ab} to be a hermitian matrix, such that $\varepsilon_{ab} = \varepsilon_{ba}$. This introduces no new undetermined multipliers into the equation, and creates a form more amenable to further derivation. Thus the Lagrangian function we will be working with is

$$
\mathcal{L}[\{\chi_a\}] = E_0[\{\chi_a\}] - \sum_{a}^{N} \sum_{b}^{N} \varepsilon_{ba} (\langle a|b \rangle - \delta_{ab}). \tag{89}
$$

The differential of this function must be set to zero as before, giving us

$$
\delta \mathcal{L} = \delta E_0 - \sum_{a}^{N} \sum_{b}^{N} \varepsilon_{ba} \delta \langle a | b \rangle = 0 \tag{90}
$$

Since we have an electronic energy in terms of spin orbitals

$$
E_0 = \sum_{a}^{N} \langle a|H|a \rangle + \frac{1}{2} \sum_{ab}^{N} \langle ab||ab \rangle
$$

\n
$$
E_0 = \sum_{a}^{N} \langle a|H|a \rangle + \frac{1}{2} \sum_{ab}^{N} [\langle ab|ab \rangle - \langle ab|ba \rangle],
$$
\n(91)

we can write the variance of the energy δE_0 as

$$
\delta E_0 = \sum_{a}^{N} (\langle \delta a | H | a \rangle + \langle a | H | \delta a \rangle)
$$

$$
+ \frac{1}{2} \sum_{ab}^{N} \left[\begin{array}{l} \langle \delta ab | ab \rangle + \langle a \delta b | ab \rangle + \langle ab | \delta ab \rangle + \langle ab | a \delta b \rangle + \\ - \langle \delta ab | ba \rangle - \langle a \delta b | ba \rangle - \langle ab | \delta ba \rangle - \langle ab | b \delta a \rangle \end{array} \right].
$$
 (92)

It takes some thought to realize that there are only two unique two electron integrals in this list, and that it an be written

$$
\delta E_0 = \sum_{a}^{N} \langle \delta a | H | a \rangle + \sum_{ab}^{N} (\langle \delta a b | ab \rangle - \langle \delta a b | ba \rangle) + \text{complex conjugate}
$$
 (93)

The other variance we need is $\varepsilon_{ba} \delta \langle a | b \rangle$, which can be expanded as

$$
\sum_{ab} \varepsilon_{ba} \langle \delta a | b \rangle + \sum_{ab} \varepsilon_{ba} \langle a | \delta b \rangle = \sum_{ab} \varepsilon_{ba} \langle \delta a | b \rangle + \sum_{ab} \varepsilon_{ab} \langle b | \delta a \rangle
$$

$$
= \sum_{ab} (\varepsilon_{ba} \langle \delta a | b \rangle + \varepsilon_{ba}^* \langle \delta a | b \rangle^*).
$$
(94)

So the whole thing boils down to one neat statement,

$$
\delta \mathcal{L} = \sum_{a} \langle \delta a | H | a \rangle + \sum_{a}^{N} \langle \delta a | b \rangle + \sum_{ab}^{N} (\langle \delta a b | a b \rangle - \varepsilon_{ba} \langle \delta a | b \rangle) + \text{complex conjugate} = 0 \quad (95)
$$

If we conveniently define a coulomb operator $J_b(1)$ and an exchange operator $K_b(1)$ as

$$
\hat{J}_b(1) = \int d\mathbf{x}_2 |\chi_b(2)|^2 r_{12}^{-1} \tag{96}
$$

$$
\hat{K}_b(1)\chi_a(1) = \int d\mathbf{x}_2 \chi_b^*(2) r_{12}^{-1} \chi_a(2) \chi_b(1), \tag{97}
$$

we can rewrite the two electron integral

$$
\langle ab||ab\rangle = \int d\mathbf{x}_1 \chi_a^*(1)(\hat{J}_b(1) - \hat{K}_b(1))\chi_a(1)
$$

=
$$
\langle a|(\hat{J}_b(1) - \hat{K}_b(1))|a\rangle.
$$
 (98)

This allows for a more compact notation to be employed in writing the variance in the lagrangian function

$$
\delta \mathcal{L} = \sum_{a} \int d\mathbf{x}_{1} \delta \chi_{a}^{*}(1) [\hat{h}(1)\chi_{a}(1) + \sum_{b} \{ (\hat{J}_{b}(1) - \hat{K}_{b}(1))\chi_{a}(1) - \varepsilon_{ba}\chi_{b}(1) \}] + \text{complex conjugate} = 0.
$$
\n(99)

Like before, the part in brackets is forced to be zero, since $\delta \chi_a(1)$ can be anything. Setting it equal to zero and rearranging to make it look like some sort of eigenvalue equation yields

$$
[\hat{h}(1) + \sum_{b} (\hat{J}_b(1) - \hat{K}_b(1))] \chi_a(1) = \sum_{b} \varepsilon_{ba} \chi_b(1)
$$
\n(100)

$$
f(1)\chi_a(1) = \sum_b \varepsilon_{ba}\chi_b(1) \tag{101}
$$

These are the glorious Hartree–Fock equations derived in general in the spin orbital basis. But wait – there's a problem. These are coupled integro-differential equations, and while they are not strictly unsolvable, they're a pain. It would be nice to at least uncouple them, so let's do that.

If we apply a unitary rotation to the full set of spin orbitals, generating a new set

$$
\chi'_a = \sum_b \chi_b U_{ba} \tag{102}
$$

where \cup is unitary, ie. $\cup^+ \equiv \cup^-$, what changes? The rotation can be written as a matrix product, if we define **A** as the matrix resembling the slater determinant for the system, ie. $(N!)^{-1/2}$ det(**A**) $= |\Psi_0\rangle$. In that case,

$$
\det(\mathbf{A}') = \det(\mathbf{A}) \det(\mathbf{U})
$$
\n(103)

However, since this matrix U is unitary, $|\det(\mathbf{U})|^2 = 1$, $\det(\mathbf{U}) = e^{i\phi}$ and the new wavefunction differs from the old by a phase factor, affecting nothing observable. How does it affect $f(1)$ and ε_{ba} ?

$$
\sum_{a} \hat{J}'_a(1) = \sum_{a} \int d\mathbf{x}_2 \chi'^*_a(2) r_{12}^{-1} \chi'_a(2)
$$
\n(104)

$$
= \sum_{a} \int d\mathbf{x}_{2} \sum_{b} U_{ba}^{*} \chi_{b}^{*}(2) r_{12}^{-1} \sum_{c} U_{ca} \chi_{c}(1) \tag{105}
$$

$$
= \sum_{bc} \left(\sum_{a} U_{ba}^{*} U_{ca} \right) \int d \mathbf{x}_{2} \chi_{b}^{*}(2) r_{12}^{-1} \chi_{c}(2) \tag{106}
$$

$$
= \sum_{bc} \delta_{bc} \int d\mathbf{x}_2 \chi_b^*(2) r_{12}^{-1} \chi_c(2) = \sum_b \hat{J}_b(1) \tag{107}
$$

$$
\sum_{a} \hat{K}'_{a}(1)\chi'_{d}(1) = \sum_{a} \int d\mathbf{x}_{2}\chi'^{*}_{a}(2)r_{12}^{-1}\chi'_{d}(2)\chi'_{a}(1)
$$
\n(108)

$$
= \sum_{a} \int d\mathbf{x}_{2} \sum_{b} U_{ba}^{*} \chi_{b}^{*}(2) r_{12}^{-1} \chi_{d}(2)' \sum_{c} U_{ca} \chi_{c}(1)
$$
 (109)

$$
= \sum_{bc} \left(\sum_{a} U_{ba}^{*} U_{ca} \right) \int d\mathbf{x}_{2} \chi_{b}^{*}(2) r_{12}^{-1} \chi_{d}(2)' \chi_{c}(1) \tag{110}
$$

$$
= \sum_{bc} \delta_{bc} \int d\mathbf{x}_2 \chi_b^*(2) r_{12}^{-1} \chi_d(2)' \chi_c(1) = \sum_b \hat{K}_b(1) \chi_d'(1) \tag{111}
$$

So $f(1)' = f(1)!$ How about ε_{ba} ? Start by realizing that ε_{ba} are matrix elements of the fock operator.

$$
\langle \chi_c | f | \chi_a \rangle = \sum_b \varepsilon_{ba} \langle \chi_c | \chi_b \rangle \tag{112}
$$

$$
= \varepsilon_{ca} \tag{113}
$$

$$
\varepsilon'_{ab} = \int d\mathbf{x}_1 \chi_a'^* f(1) \chi_b'(1) \tag{114}
$$

$$
= \sum_{cd} U_{ca}^* U_{db} \varepsilon_{cd} \tag{115}
$$

$$
= \mathbf{U}^{\dagger} \varepsilon \mathbf{U} \tag{116}
$$

This last result can be written as a martix product as well, and it is seen that this is now a unitary transformation to the matrix ε . We are free to choose U to be whatever we please, and if we choose it to make ε diagonal, we can rewrite the Hartree–Fock equations as

$$
f|\chi_a\rangle = \varepsilon_a|\chi_a\rangle. \tag{117}
$$

When this is done, the resulting spin orbitals are termed the Hartree-Fock canonical orbitals. Read section 3.3 in Szabo and Ostlund for various fun things to do with the Hartree–Fock equations.

The problem still remains, though. These are integro-differential equations, which computers (and omputer programmers) balk at. That is why Roothaan is a HERO! Through his results, we can transform these into a set of matrix formulated algebraic equations that computers and programmers dig. The general case is too troublesome for now, so let's limit ourselves to closed shell, RHF orbitals. To take advantage of the simplifications this can afford, we need to return to the spatial orbital basis. We've derived the spin orbital based Fo
k operator

$$
f(\mathbf{x}_1) = \hat{h}(\mathbf{x}_1) + \sum_{a}^{N} (\hat{J}_a(\mathbf{x}_1) - \hat{K}_a(\mathbf{x}_1)).
$$
\n(118)

Without further ado, we'll introduce the spatial orbital based Fock operator and be done with it.

$$
f(\mathbf{r}_1) = \hat{h}(\mathbf{r}_1) + \sum_{a}^{N/2} (2\hat{J}_a(\mathbf{r}_1) - \hat{K}_a(\mathbf{r}_1))
$$
(119)

$$
\hat{J}_b(\mathbf{r}_1) = \int d\mathbf{r}_2 |\psi_b(\mathbf{r}_2)|^2 r_{12}^{-1} \tag{120}
$$

$$
\hat{K}_b(\mathbf{r}_1)\psi_a(\mathbf{r}_1) = \int d\mathbf{r}_2 \psi_b^*(\mathbf{r}_2) r_{12}^{-1} \psi_a(\mathbf{r}_2) \psi_b(1) \qquad (121)
$$

With the properties

$$
f(\mathbf{r}_1)\psi_j(\mathbf{r}_1) = \varepsilon_j \psi_j(\mathbf{r}_1) \tag{122}
$$

$$
\varepsilon_i = h_{ii} + \sum_{b}^{N/2} 2J_{ib} - K_{ib} \tag{123}
$$

Now we introduce a basis set expansion to bring the HF integro-differential equations to soluable algebraic equations. Letting $\psi_i = \sum_{\mu} C^i_{\mu} \phi_{\mu}$,

$$
f(1) \sum_{\nu} C_{\nu}^{i} \phi_{\nu} = \varepsilon_{i} \sum_{\nu} C_{\nu}^{i} \phi_{\nu}
$$
 (124)

Multiplying by $\varphi_{\mu}(1)$ and integrating over electron 1 gives

$$
\sum_{\nu} C_{\nu}^{i} \int d\mathbf{r}_{1} \phi_{\mu}^{*}(1) f(1) \phi_{\nu}(1) = \varepsilon_{i} \sum_{\nu} C_{\nu}^{i} \int d\mathbf{r}_{1} \phi_{\mu}^{*}(1) \phi_{\nu}(1).
$$
 (125)

Identifying the integrals as matrix elements of the fock operator and the unit operator (overlap) respe
tively,

$$
\sum_{\nu} F_{\mu\nu} C_{\nu}^{i} = \varepsilon_{i} \sum_{\nu} S_{\mu\nu} C_{\nu}^{i}
$$
\n(126)

Using the fact that ε is diagonal, this can be written as the matrix product

$$
\mathbf{FC} = \mathbf{SC}_{\varepsilon} \tag{127}
$$

So what is **F**, this so called Fock Matrix? We've defined $F_{\mu\nu}$ before as the matrix element of the one electron fock operator, $f(1)$ in equation 119. Writing out this integral and expanding,

$$
F_{\mu\nu} = \int d\mathbf{r}_{1} \phi_{\mu}^{*}(1) f(1) \phi_{\nu}(1)
$$

\n
$$
= \int d\mathbf{r}_{1} \phi_{\mu}^{*}(1) [h(1) + \sum_{a} (2\hat{J}_{a}(1) - \hat{K}_{a}(1))] \phi_{\nu}(1)
$$

\n
$$
= \int d\mathbf{r}_{1} \phi_{\mu}^{*}(1) h(1) \phi_{\nu}(1) + \sum_{a} [2 \int d\mathbf{r}_{1} \phi_{\mu}^{*}(1) \hat{J}_{a}(1) \phi_{\nu}(1) - \int d\mathbf{r}_{1} \phi_{\mu}^{*}(1) \hat{K}_{a}(1) \phi_{\nu}(1)] \quad (128)
$$

\n
$$
= h_{\mu\nu} + \sum_{a} 2(\mu\nu|aa) - (\mu a|a\nu)
$$

\n
$$
= h_{\mu\nu} + \sum_{a} [\sum_{\rho\sigma} C_{\rho}^{a} C_{\sigma}^{a} 2(\mu\nu|\rho\sigma) - (\mu\rho|\sigma\nu)]
$$

\n
$$
= h_{\mu\nu} + \sum_{\rho\sigma} (\sum_{a} C_{\rho}^{a} C_{\sigma}^{a}) [2(\mu\nu|\rho\sigma) - (\mu\rho|\sigma\nu)]
$$

\n
$$
= h_{\mu\nu} + \sum_{\rho\sigma} \mathbf{D}_{\rho\sigma} [2(\mu\nu|\rho\sigma) - (\mu\rho|\sigma\nu)] \qquad (129)
$$

This is a quantity which can be easily constructed given a set of molecular orbitals (the coefficients C^a_μ) and a precalculated set of atomic orbital integrals. At this point, the Hartree–Fock equations have been reduced to a matrix eigenvector problem, $FC = SC\varepsilon$, but not in a computationally convenient form. Following the analysis leading to equation 84, we first define the transformed Fo
k matrix as

$$
\mathbf{F}^t = \mathbf{S}^{-1/2} \mathbf{F} \mathbf{S}^{-1/2}.
$$
 (130)

We an then take

$$
\mathbf{F}^t(\mathbf{S}^{1/2}\mathbf{C}) = (\mathbf{S}^{1/2}\mathbf{C})\varepsilon\tag{131}
$$

to be equivalent to equation 127. Since ε is diagonal by choice, diagonalizing the transformed Fock matrix gives a set of transformed molecular orbital coefficients, $C^{\dagger} = S^{1/2}C$. The matrix $S^{1/2}C$ can be back transformed to give the true MO coefficient matrix, C. The density matrix for these coefficients is formed by the product $D = C^{\dagger}C$, and can subsequently be used to construct a new fock matrix via equation 129. Since the overlap matrix S does not depend on the MO coefficients, the same unitary transformation can be applied to the new fock matrix to give a new transformed fock matrix. This can be diagonalized to produce new MO coefficients, and the process repeated until onvergan
e. As an initial guess for the fo
k matrix, one generally uses the ore hamiltonian, ignoring all the two electron integrals.

$$
F_{\mu\nu}^{1} = H_{\mu\nu} = \int \phi_{\mu}^{*} (-\nabla_{i}^{2} + \sum_{A} \frac{Z_{A}}{r_{iA}}) \phi_{\nu}
$$
(132)

From the core hamiltonian, an initial C is obtained and a much improved Fock matrix can be built including the two electron integrals. And that's about it for the Hartree-Fock Self Consistent Field Method. These last few pages will be the most important when you get around to programming the closed shell SCF method for the specific case of water, as you will be given the integrals in a file, and you can begin the process by building the core hamiltonian as described above.