

Hvězdná noc, 1889

Struktura a kinematika galaxíí

Bruno Jungwiert



II. Hubbleova a de Vaucouleursova klasifikace
II.a. Prstence v diskových galaxiích

1920-1924: Spiral nebulae (galaxies) and the „Great Debate” questions:

- 1) Are they made of stars or gas?
- 2) Are they inside or outside of our own Galaxy? If outside, how far?
- 3) How big is our Galaxy? 30,000 l.y.? 300,000 l.y.?

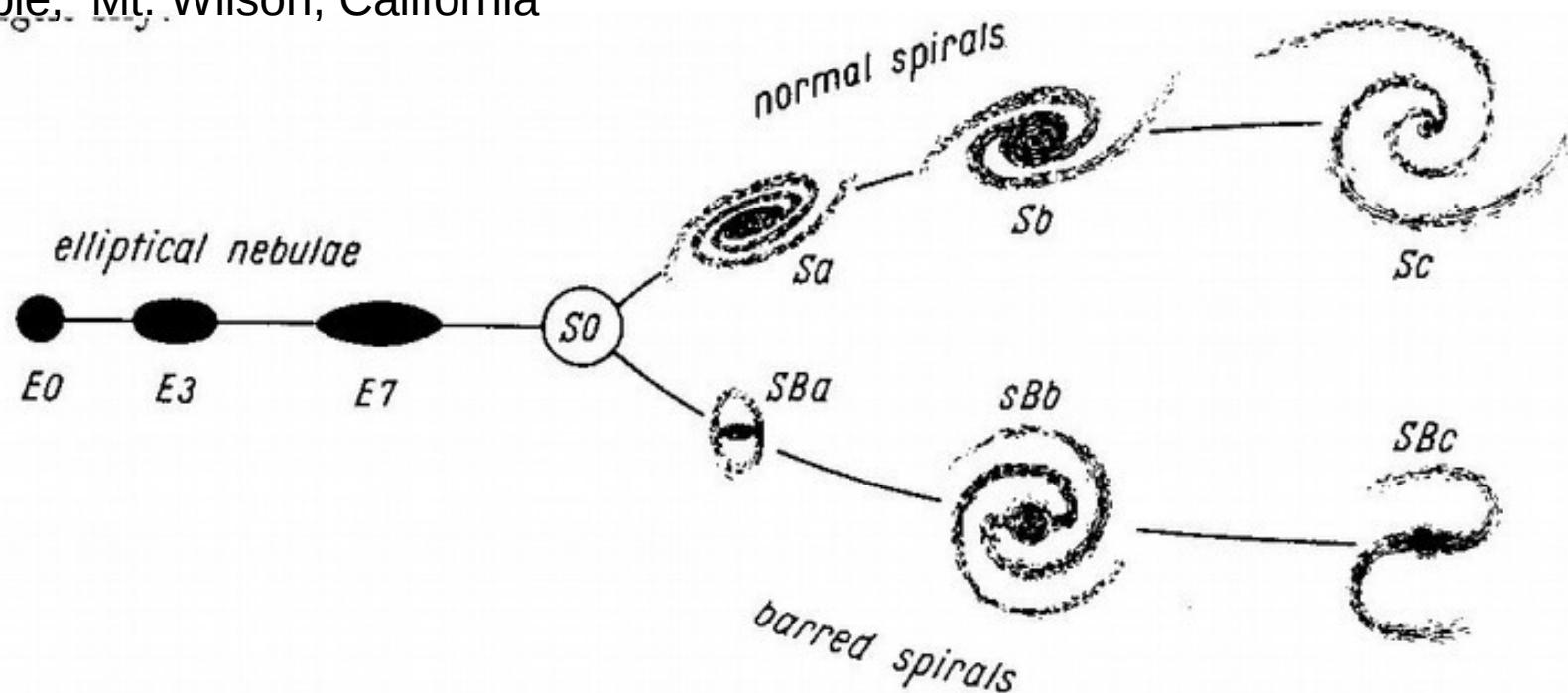
1924: Two important discoveries are made by Edwin Hubble:

- 1) some individual stars (**Cepheids**) resolved in nearby “spiral nebulae”
- 2) distances to these stars, and thus to their host galaxies, estimated
=> conclusions: “spiral nebulae” are made of stars; they are independent *stellar islands* far beyond the outskirts of our own Galaxy

1926-1936: The Hubble Sequence (The Hubble classification of galaxies =

the first classification of galaxy morphologies)

Edwin P. Hubble, Mt. Wilson, California



from: G. de Vaucouleurs, Classification and Morphology of External Galaxies, Handbuch der Physik, 1959, Vol. 53, pp. 275-210

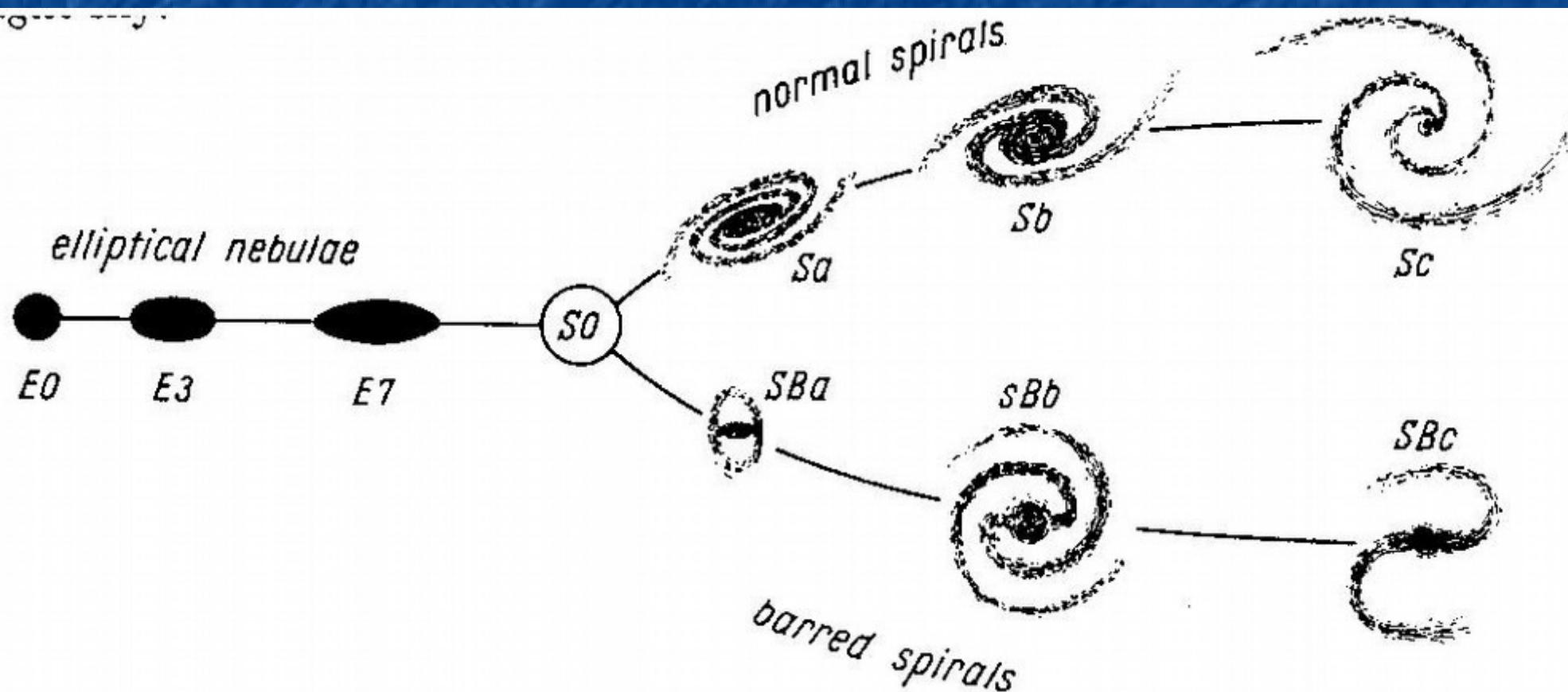
<http://nedwww.ipac.caltech.edu/level5/Dev/frames.html>

The Hubble tuning fork (1925, Mt. Wilson)

Elliptical (E), lenticular (S0), normal spiral (S), barred spiral (SB) and irregular galaxies (I)

E sub-types: E0 – E7, $n = 10(1 - b/a)$

S and SB sub-types: Sa, Sb, Sc and SBa, SBb, SBc

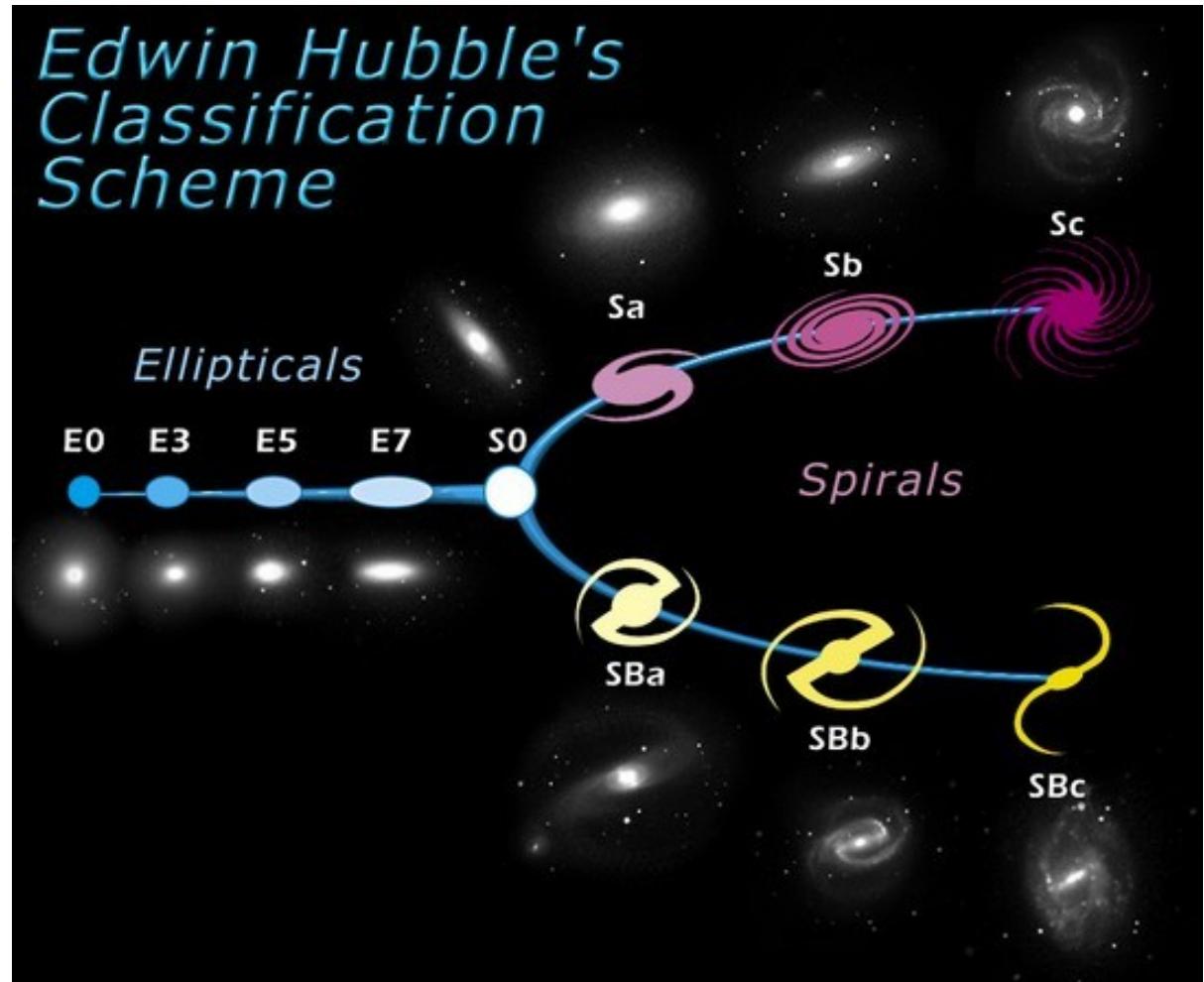


from: G. de Vaucouleurs, Classification and Morphology of External Galaxies, Handbuch der Physik, 1959, Vol. 53, pp. 275-310

<http://nedwww.ipac.caltech.edu/level5/Dev/frames.html>



Edwin P. Hubble
(1889 - 1953)



Galaxies are classified by their morphology as:
elliptical galaxies (ellipticals) (E)
lenticular galaxies (lenticulars) (S0)
spiral galaxies (spirals)
- divided into: normal spirals (S)
barred spirals (SB)

Note: The Hubble classification does not include **dwarf galaxies** that dominate the galaxy counts (they are subdivided into dwarf irregular, dwarf elliptical, dwarf spheroidal and ultra-faint dwarf galaxies).



Hooker telescope (100-inch mirror) Mt. Wilson, California

Fundamental discoveries:

- nature of spiral nebulae: **galaxies**
- speeds of galaxies: **expansion of the Universe**



Revised Hubble classification (de Vaucouleurs, 1959)

HUBBLE-DE VAUCOULEURS DIAGRAM

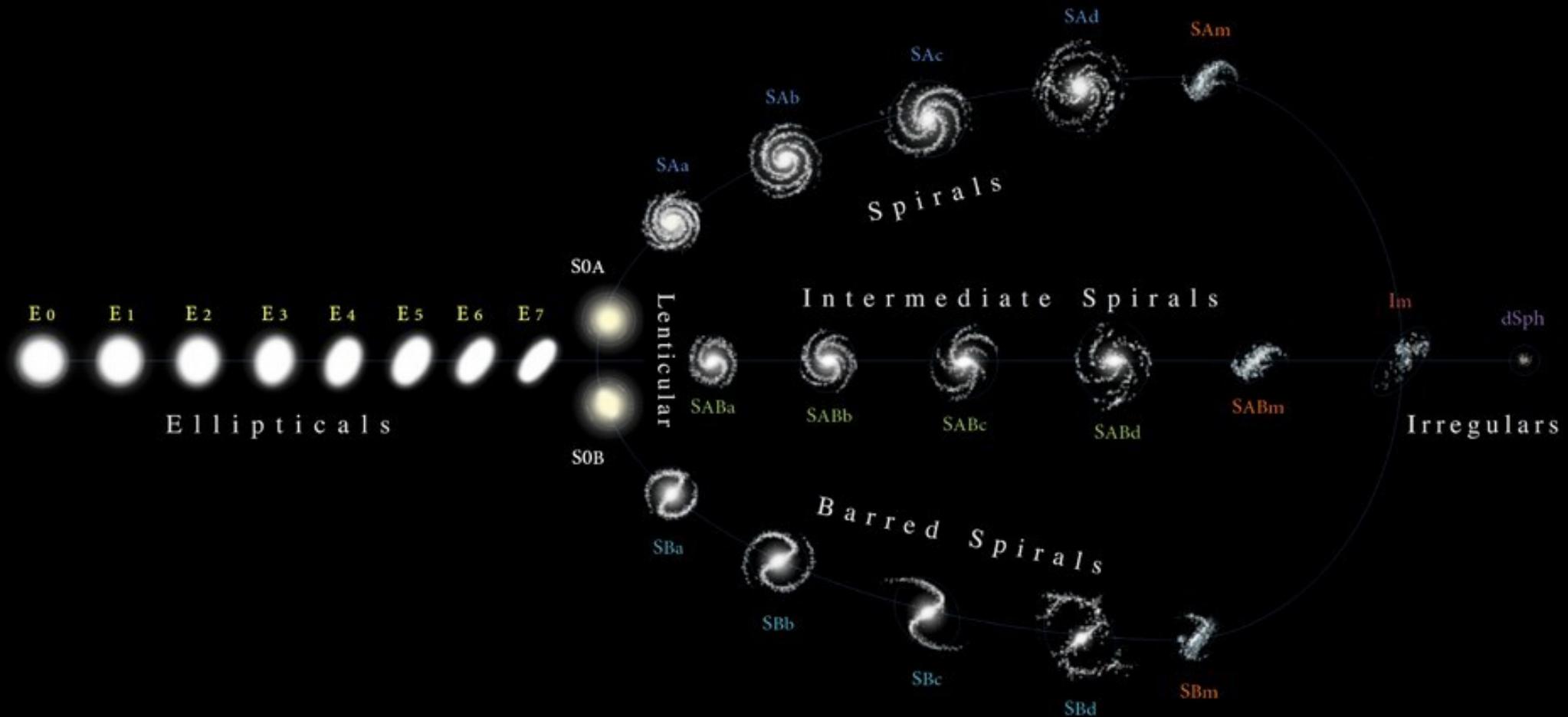


Image credit: By Antonio Ciccolella / M. De Leo - <https://en.wikipedia.org/wiki/File:Hubble-Vaucouleurs.png>, CC BY 3.0, <https://commons.wikimedia.org/w/index.php?curid=50260841>

AAT 60

elliptical galaxies,
normal spiral galaxies,
barred spiral galaxies

face-on views
edge-views



Spiral Galaxy Messier 83
(FORS / VLT)

ESO PR Photo 24b/05 (August 10, 2005)



Spiral Galaxy NGC 4565
(FORS / VLT)

ESO PR Photo 24a/05 (August 10, 2005)



AAT 55



The Sombrero Galaxy (VLT ANTU + FORS1)

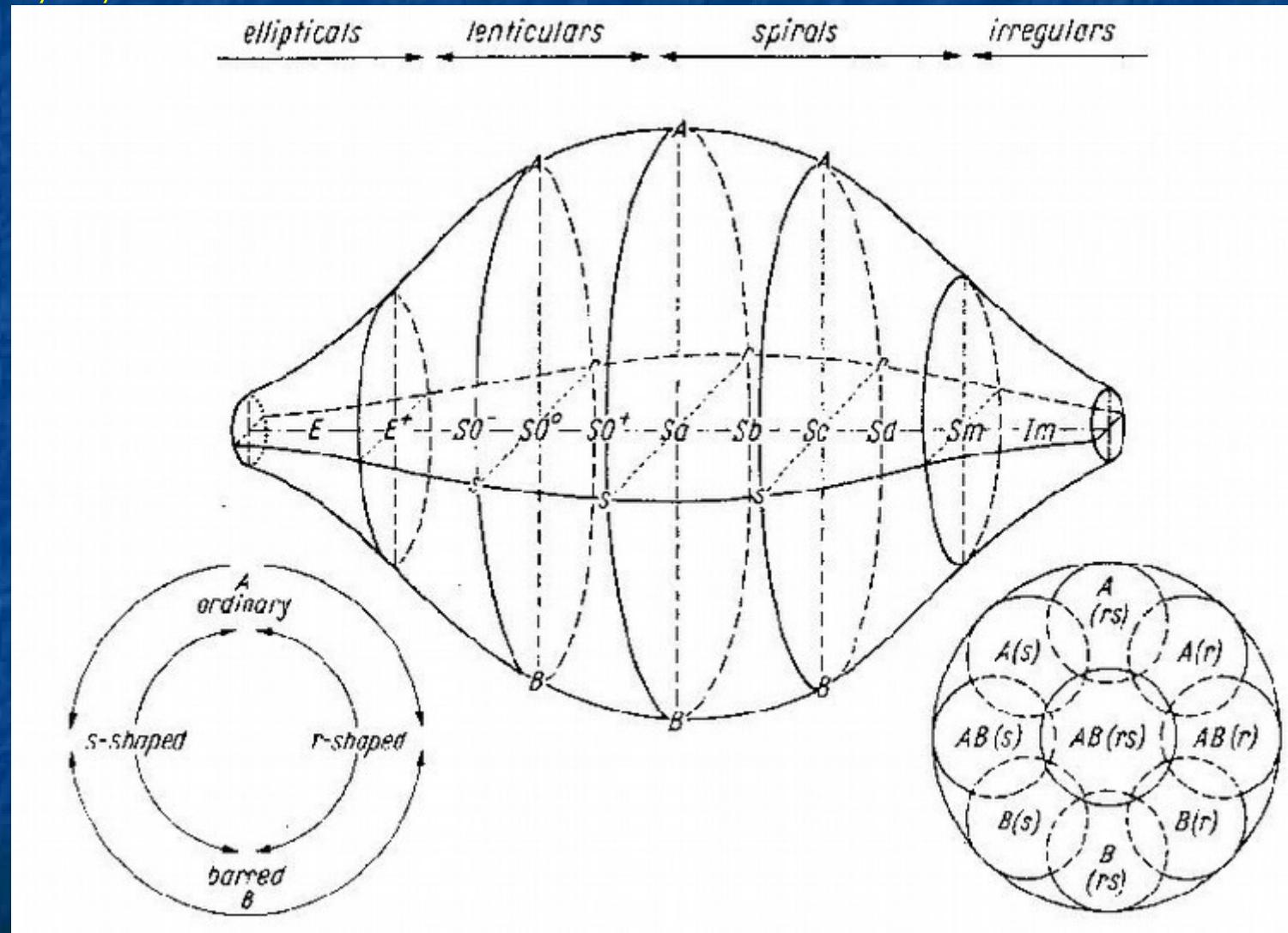
ESO PR Photo 07a/09 (22 February 2009)

© European Southern Observatory



The de Vaucouleurs classification (1959, 1964)

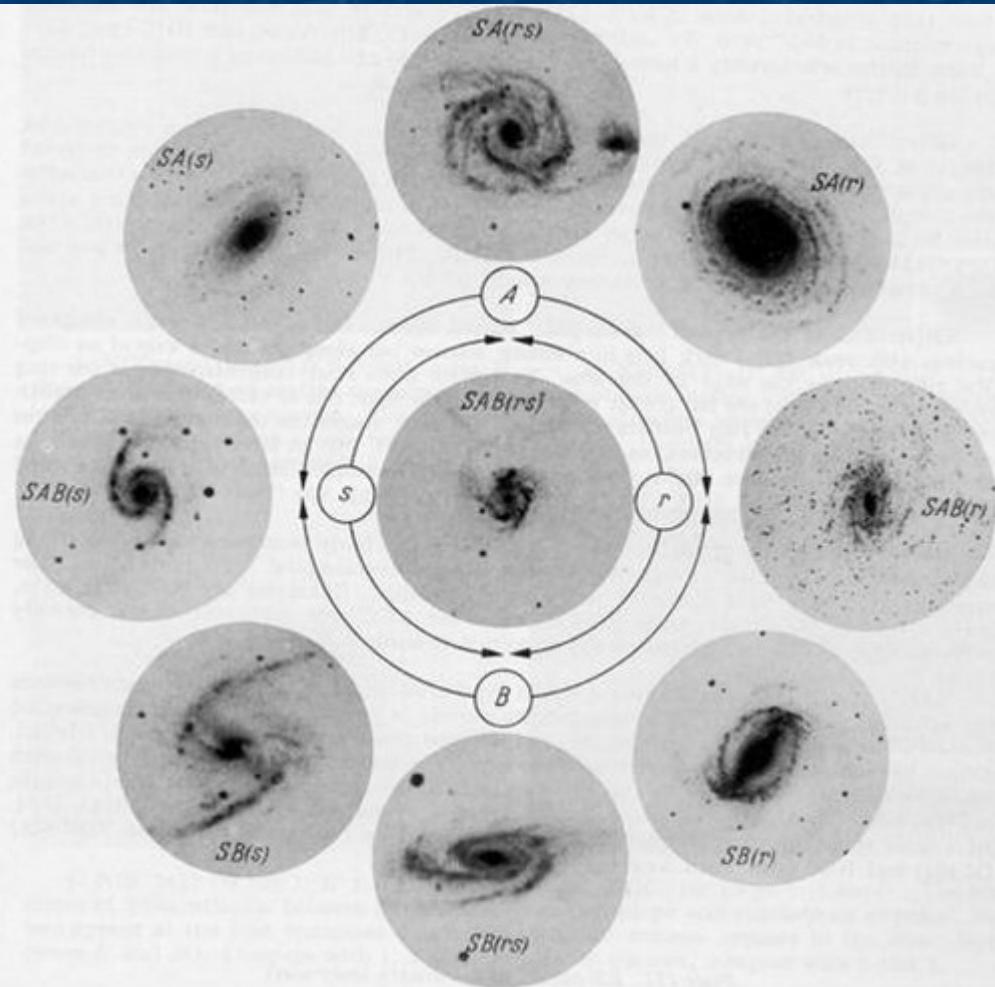
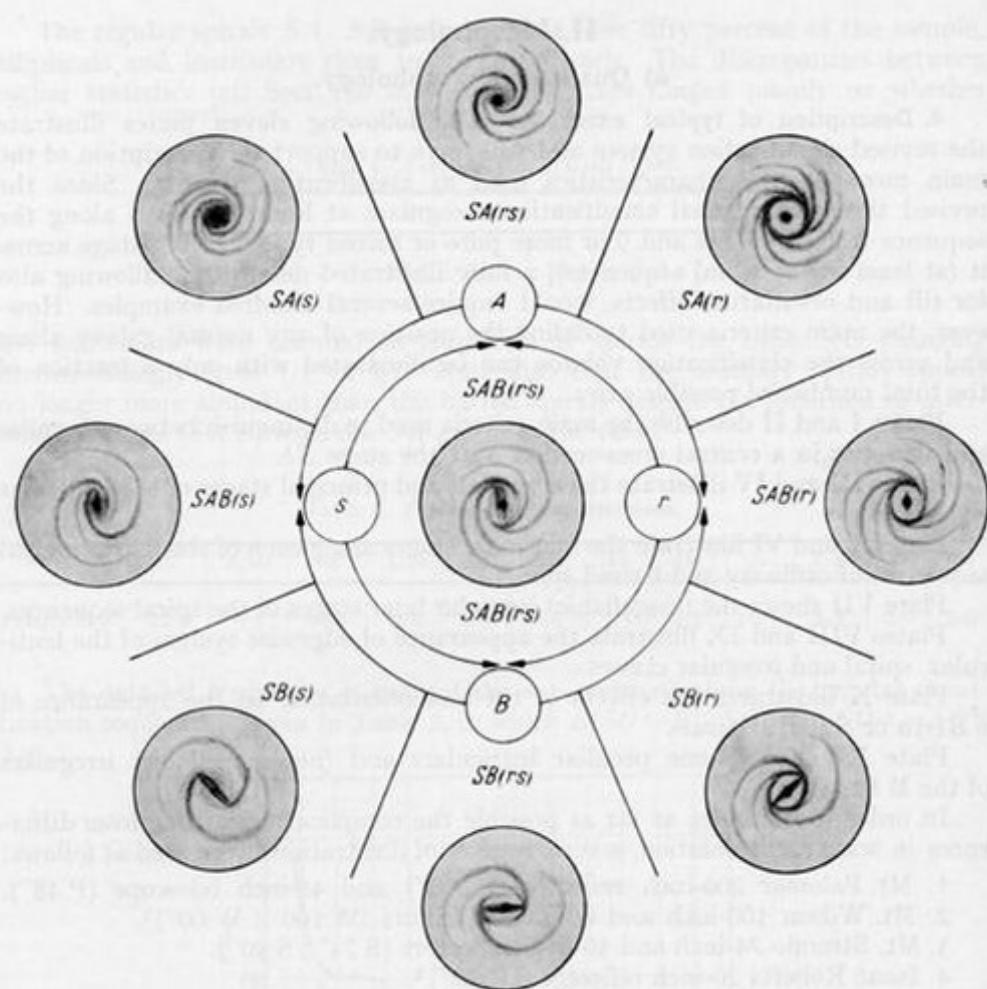
- classes E, S0, S, Im
- families A, AB, B
- varieties s, rs, r



from: G. de Vaucouleurs, Classification and Morphology of External Galaxies, Handbuch der Physik, 1959, Vol. 53, pp. 275-310

<http://nedwww.ipac.caltech.edu/level5/Dev/frames.html>

De Vaucouleurs (1959)



from: G. de Vaucouleurs, Classification and Morphology of External Galaxies, Handbuch der Physik, 1959, Vol. 53, pp. 275-310

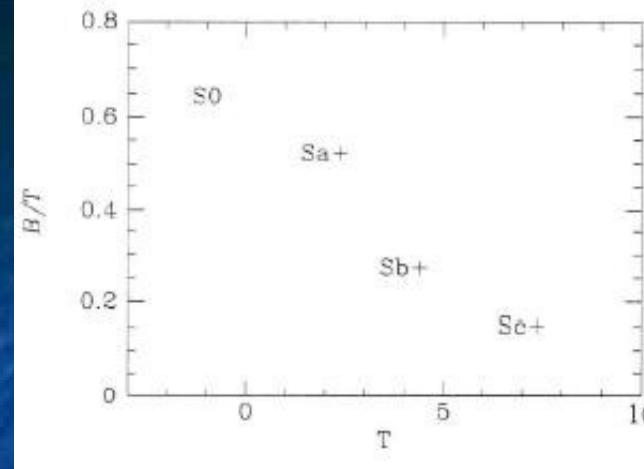
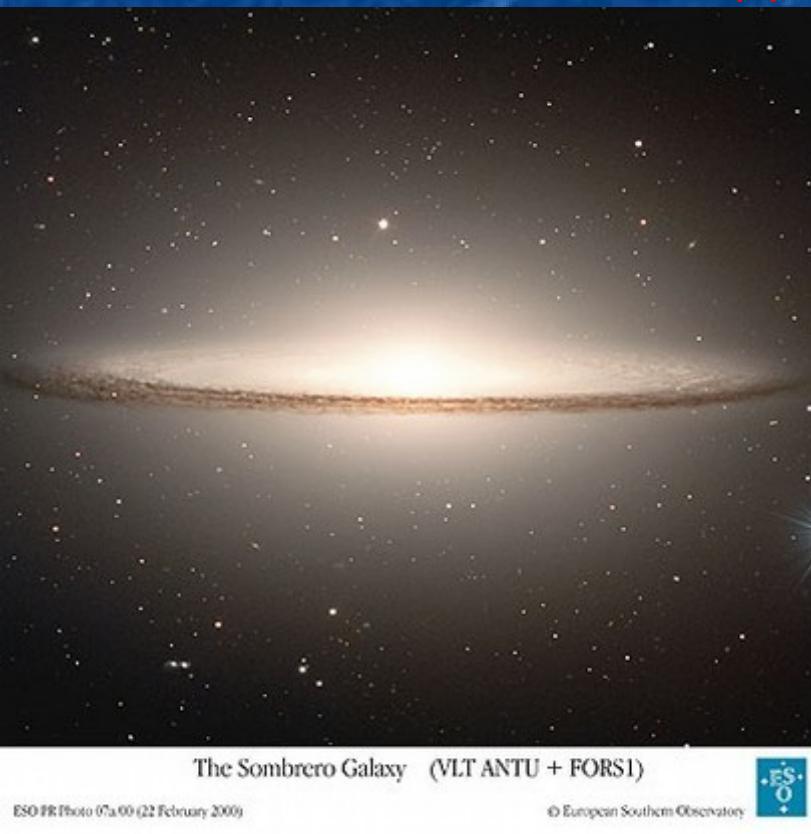
<http://nedwww.ipac.caltech.edu/level5/Dev/frames.html>

Bulges

M 104 SA(s)a
NGC 4565 SA(s)b?

B/D – bulge-to-disk ratio
 $B/T = B/(B+D)$ – bulge-to-total luminosity ratio

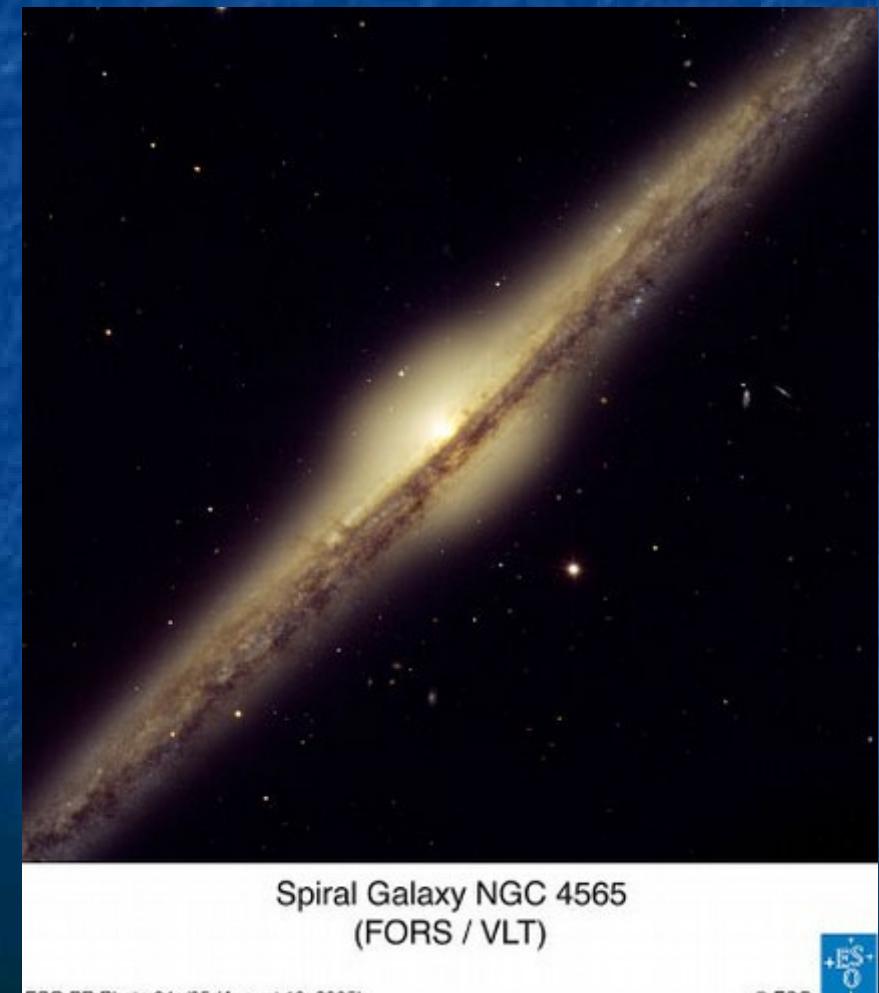
SA(s)a



(zdroj: Binney & Merrifield,
Galactic Astronomy, kapitola IV)

Figure 4.51 The later the Hubble stage T of a disk galaxy, the smaller is its bulge fraction B/T . The plotted values are means. Individual values scatter significantly about these means. [From data published in Kent (1985)]

SA(s)b



Bars (příčky)

> 2/3 diskových galaxií mají příčku viditelnou v optickém oboru

typ SA: galaxie bez příčky ~1/3

typ SAB: galaxie se „slabou“ příčkou ~1/3

typ SB: galaxie se „silnou“ příčkou ~1/3

v blízké infračervené oblasti je frekvence příček > 80%

SB(r'l)b

SAB(s)c



Spiral Galaxy Messier 83
(FORS / VLT)



SAB(rs)cd



Spiral Galaxy NGC 7424
(VLT MELIPAL + VIMOS)



Spiral Galaxy NGC 1097
(VLT MELIPAL + VIMOS)

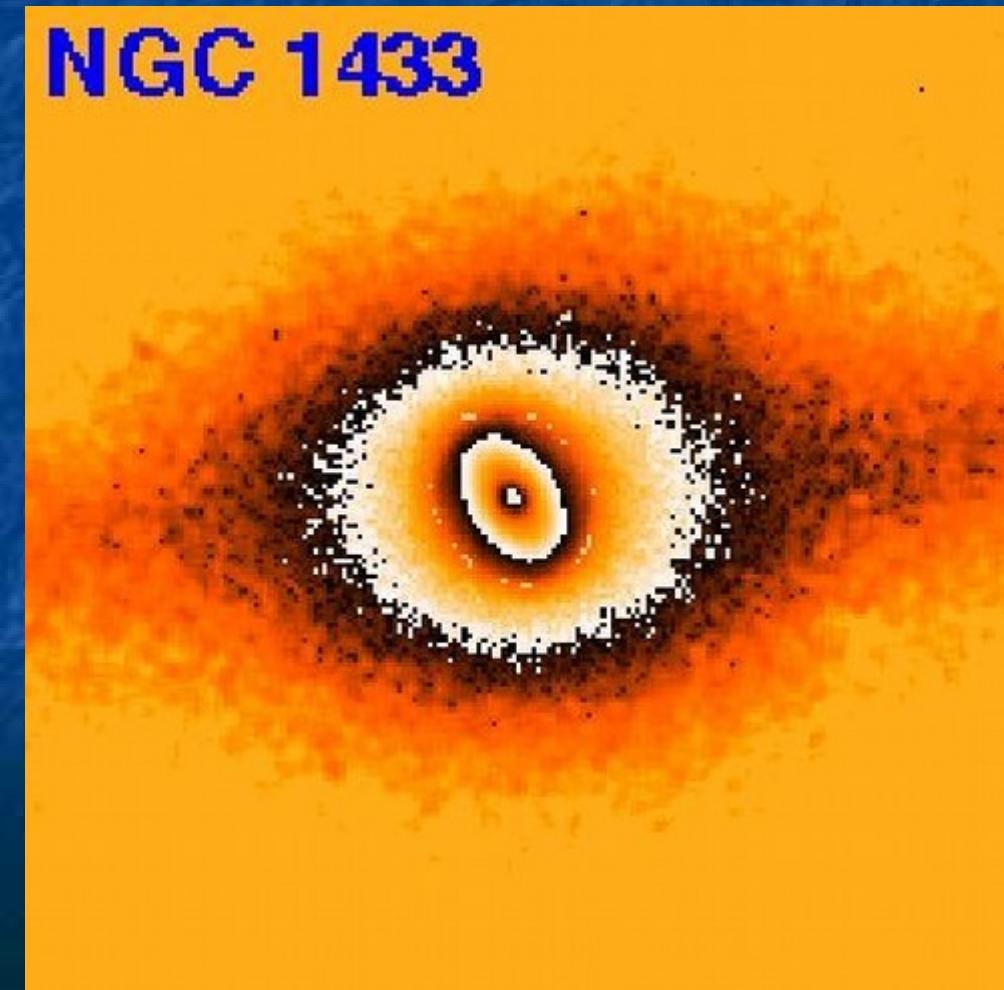


Dvojpříčky (double bars, bars-in-bars)
Nukleární příčky (nuclear bars, baby bars)
Double bar in NGC 1433 (R'_{1_1} SB(rs)ab)

(B band, Buta & Combes 1996)



(H band, JCA, 1997)



Prstence (rings) a pseudo-prstence (pseudo-rings) v diskových galaxiích

označení:

	prstence	pseudo-prstence
- vnější (outer):	R, R_1, R_2	R', R'_1, R'_2
- vnitřní (inner):	r	r_s (nebo r')
- nukleární:	nr	nr'

Příklady vnějších a vnitřních prstenců

NGC 7020 (Buta & Combes 1996)
 $(R)SA(r)0^+$



The Colossal Cosmic Eye NGC 1350
(FORS/VLT)

ESO PR Photo 31a/05 (September 27, 2005)

© ESO

$(R'_1)SB(r)ab$ (Sy)

$(R'_2)SAB(r)b$



Spiral Galaxy ESO 269-57
(VLT ANTU + FORSI)

© European Southern Observatory

Nukleární prstenec v NGC 1097

SB(r'l, nr)b



Spiral Galaxy NGC 1097
(VLT MELIPAL + VIMOS)

ESO PR Photo 35d/04 (22 December 2004)



© European Southern Observatory



The Centre of the Active Galaxy NGC 1097
(NACO/VLT)

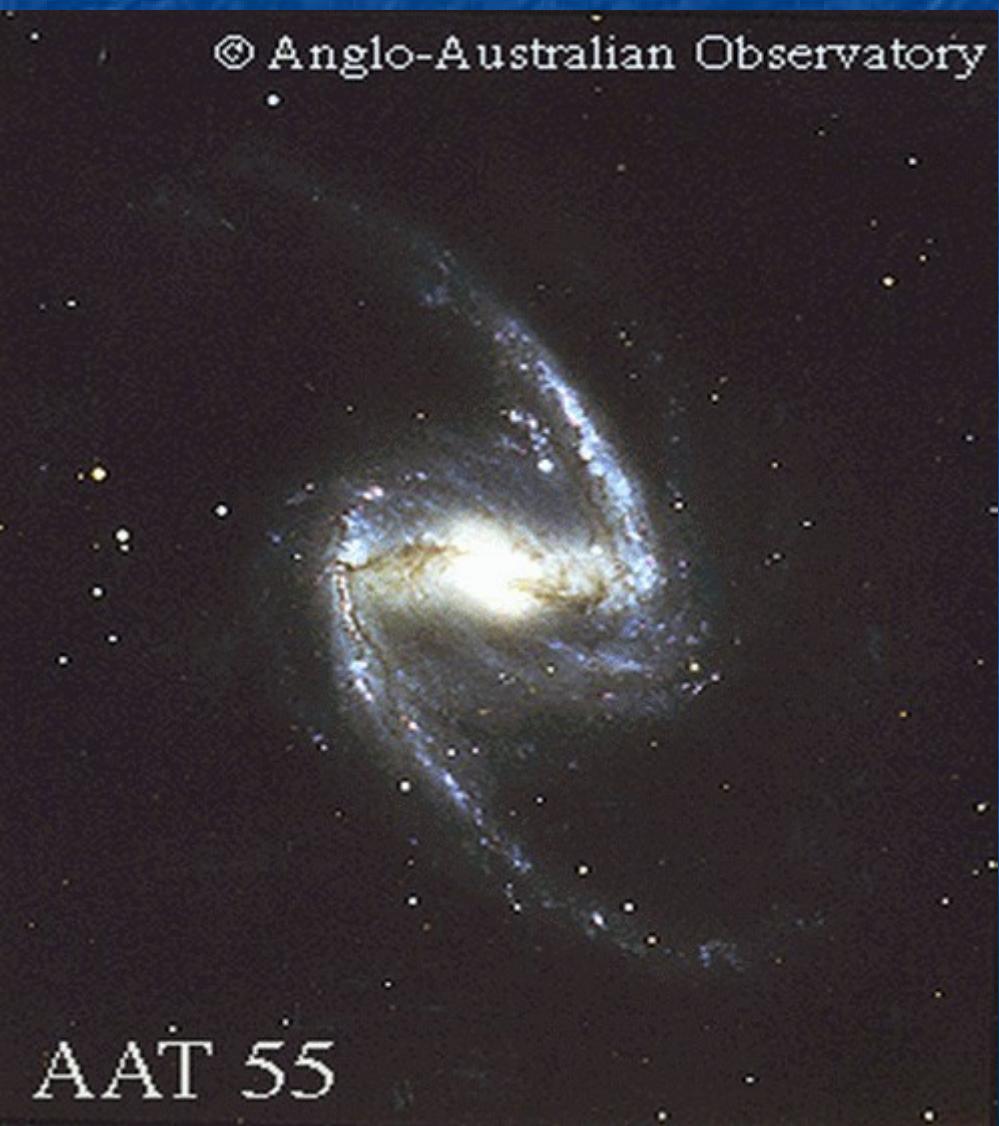
ESO PR Photo 33a/05 (October 17, 2005)

© ESO



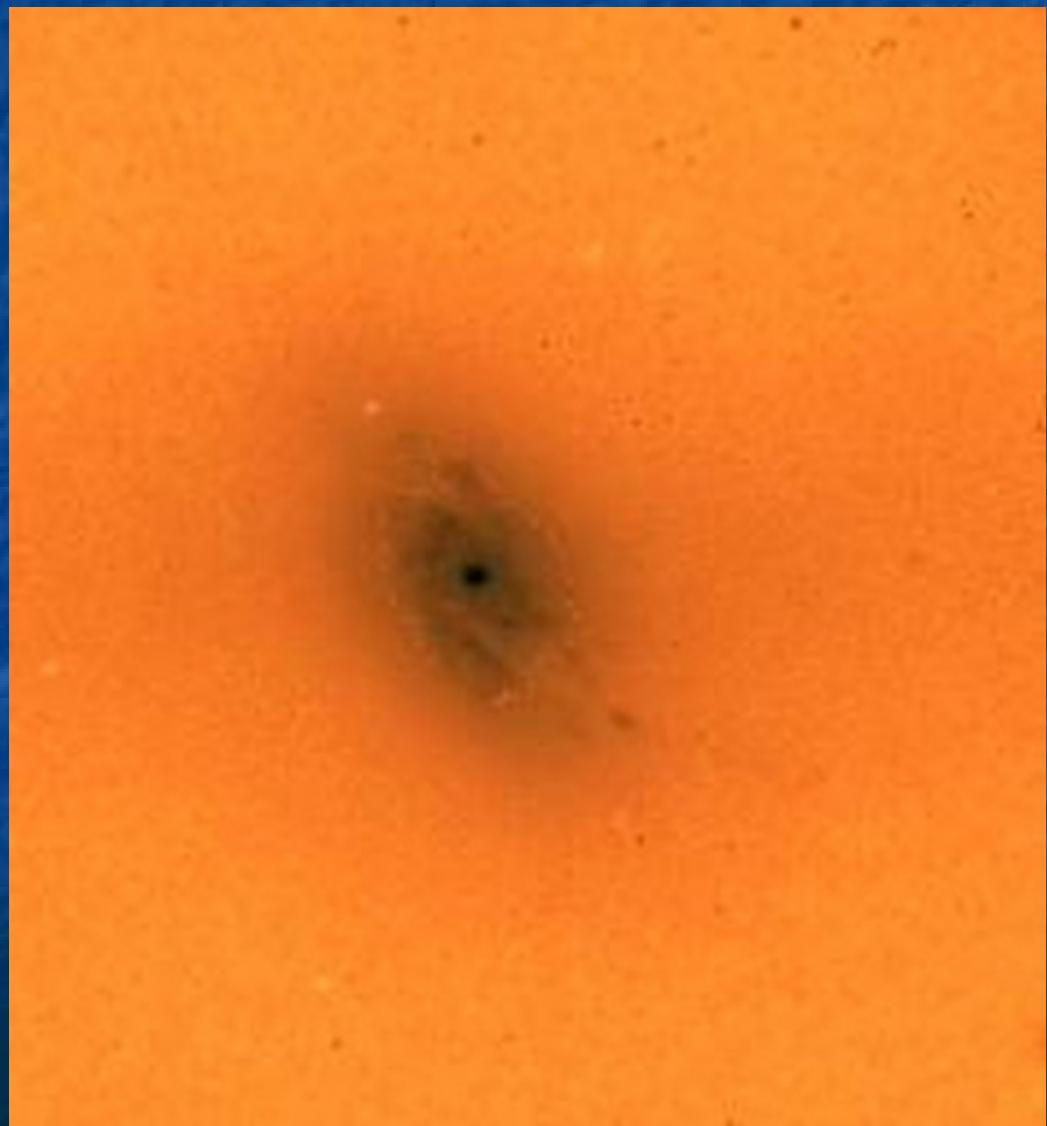
Nukleární spirály

NGC 1365, SB(s)b
optical image



near-IR image (H-band)

(different scale; from JCA, 1997)



Normal (unbarred) spiral galaxies (Hubble type SA)



M81 (Bode's galaxy) (SAab)



NGC 891 (SAb)



M64 (Black eye) (SAab)



M94 (SAab)



M104 (Sombrero) (SAa)

Normal (unbarred) spiral galaxies (Hubble type SA)



M88 (SAb)



M51 (Whirlpool galaxy) (SAbc)



M74 (Phantom) (SAc)

NGC 4565 (Needle)



M63 (Sunflower galaxy) (SAbc)



M33 (Triangulum) (SAcd)



Normal (unbarred) spiral galaxies (Hubble type SA)



M81 (Bode's galaxy) (SAab)
M94 (SAab)



NGC 891 (SAb)



M64 (Black eye) (SAab)
M104 (Sombrero) (SAa)



Weakly barred spiral galaxies (Hubble type SAB)



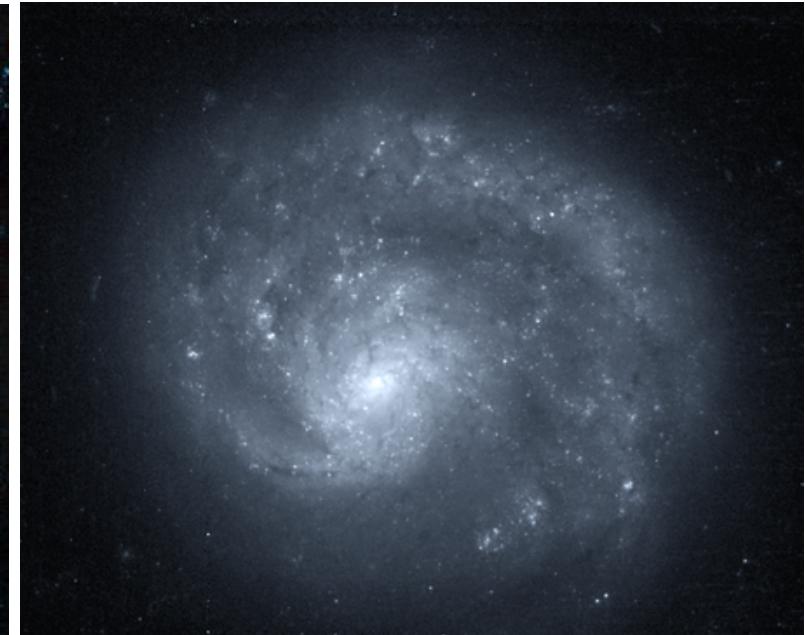
M90 (SABab)

M96 (SABab)



M61 (SABbc)

M101 (Pinwheel) (SABcd)



NGC 4625 (SABm)

M83 (Southern pinwheel) (SABc)



Strongly barred spiral galaxies (Hubble type SB)



M91 (SBb)



NGC 1365 (SBb)



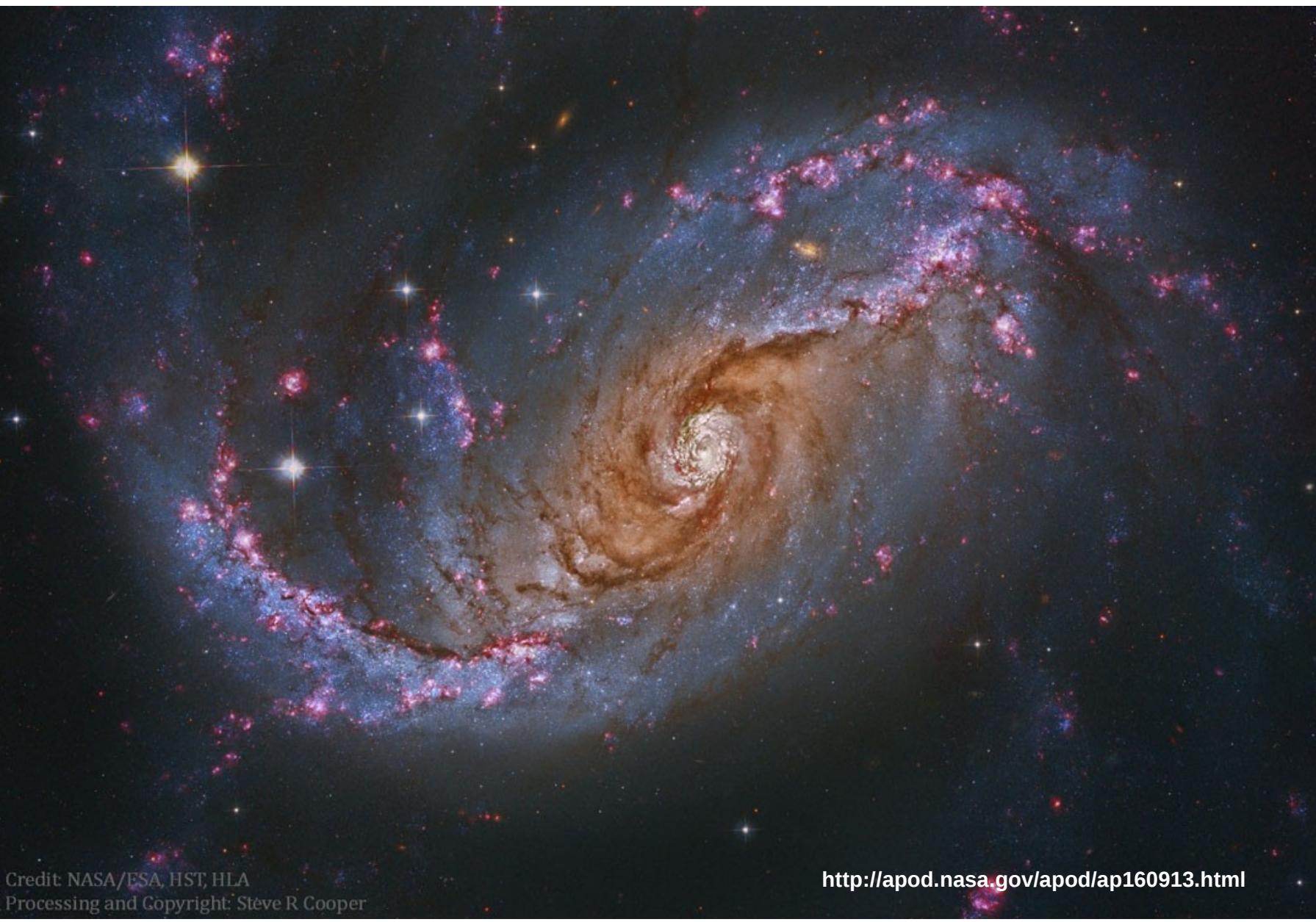
NGC 1097 (SBb)

NGC 1300 (SBbc)



NGC 2903 (SBd)





Credit: NASA/ESA, HST, HLA
Processing and Copyright: Steve R Cooper

<http://apod.nasa.gov/apod/ap160913.html>

NGC 1672: Barred Spiral Galaxy from Hubble

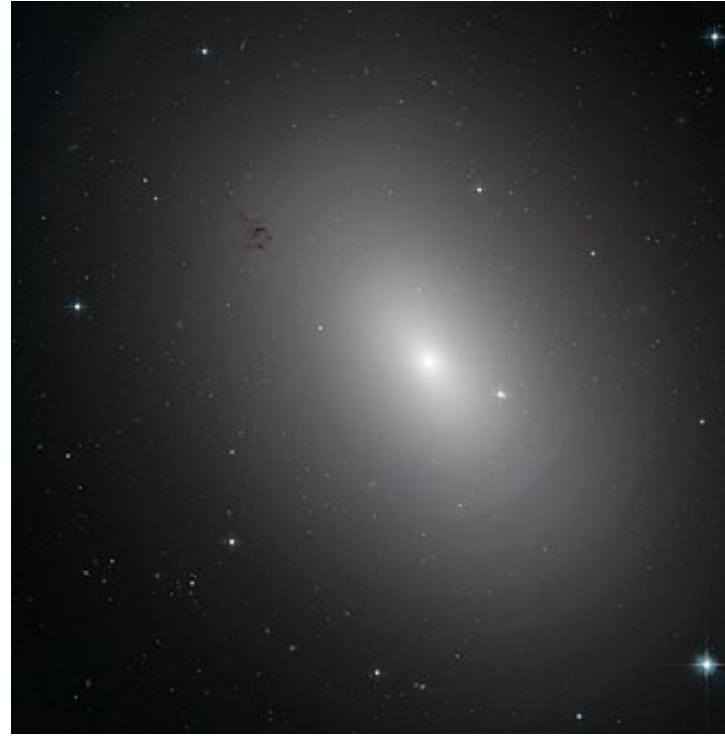
Image Credit: Hubble Legacy Archive, NASA, ESA; Processing & Copyright: Steve Cooper

Explanation: Many spiral galaxies have bars across their centers. Even our own Milky Way Galaxy is thought to have a modest central bar. Prominently barred spiral galaxy NGC 1672, featured here, was captured in spectacular detail in an image taken by the orbiting Hubble Space Telescope. Visible are dark filamentary dust lanes, young clusters of bright blue stars, red emission nebulae of glowing hydrogen gas, a long bright bar of stars across the center, and a bright active nucleus that likely houses a supermassive black hole. Light takes about 60 million years to reach us from NGC 1672, which spans about 75,000 light years across. NGC 1672, which appears toward the constellation of the Dorado (Dolphinfish), is being studied to find out how a spiral bar contributes to star formation in a galaxy's central regions.

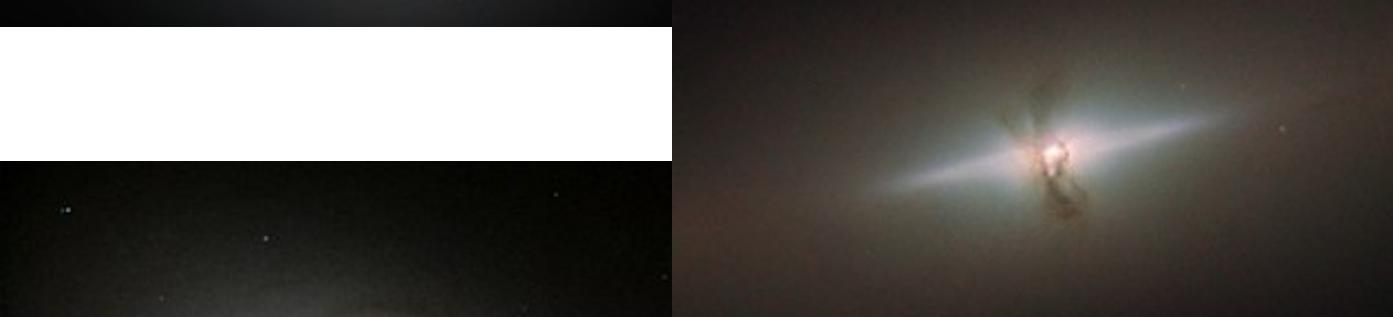
Elliptical galaxies (Hubble type E)

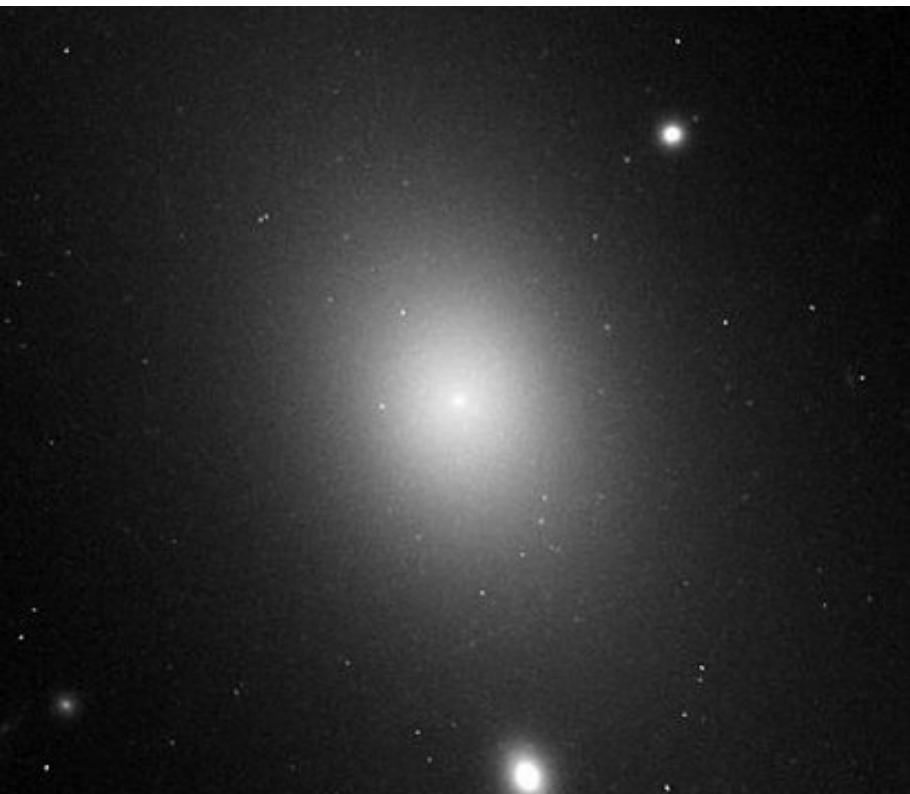


AAT 60



Lenticular galaxies (Hubble type S0)





IC 1101 – the largest known galaxy
supergiant elliptical galaxy (Hubble type E)
size: 5 million l.y. across
(25 times the Milky Way disk)
distance: 1 billion l.y. (redshift: 0.08)
number of stars: 100 trillion (10^{14})
(500 x more than the Milky Way Galaxy)